

# Azimuthal anisotropies in coherent $\rho^0$ production

Guillermo Contreras

Czech Technical University in Prague

Děčín  
September 15, 2023



## 6. miniworkshop difrakce a ultraperiferálních srážek

Analysis performed by **Andrea Riffero**

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.... QM interference .... and spin degrees of freedom ...

# Further information

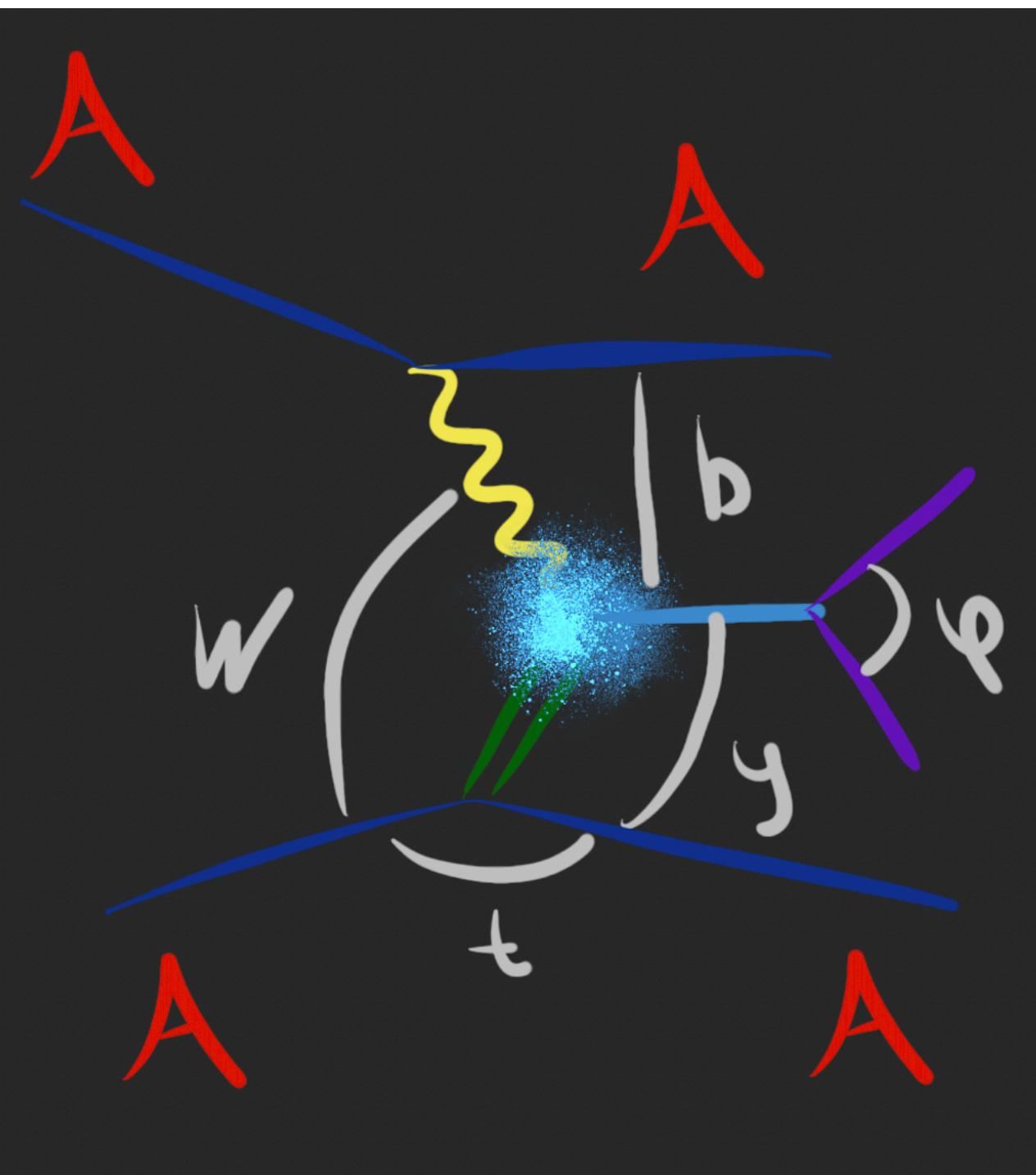
AN: <https://alice-notes.web.cern.ch/node/1387>

PF: [https://indico.cern.ch/event/1310105/contributions/5527292/attachments/2694526/4681497/azimuthAnis\\_QM\\_preliminary\\_Riffero.pdf](https://indico.cern.ch/event/1310105/contributions/5527292/attachments/2694526/4681497/azimuthAnis_QM_preliminary_Riffero.pdf)

QM poster: <https://alice-conferences.web.cern.ch/node/66110>

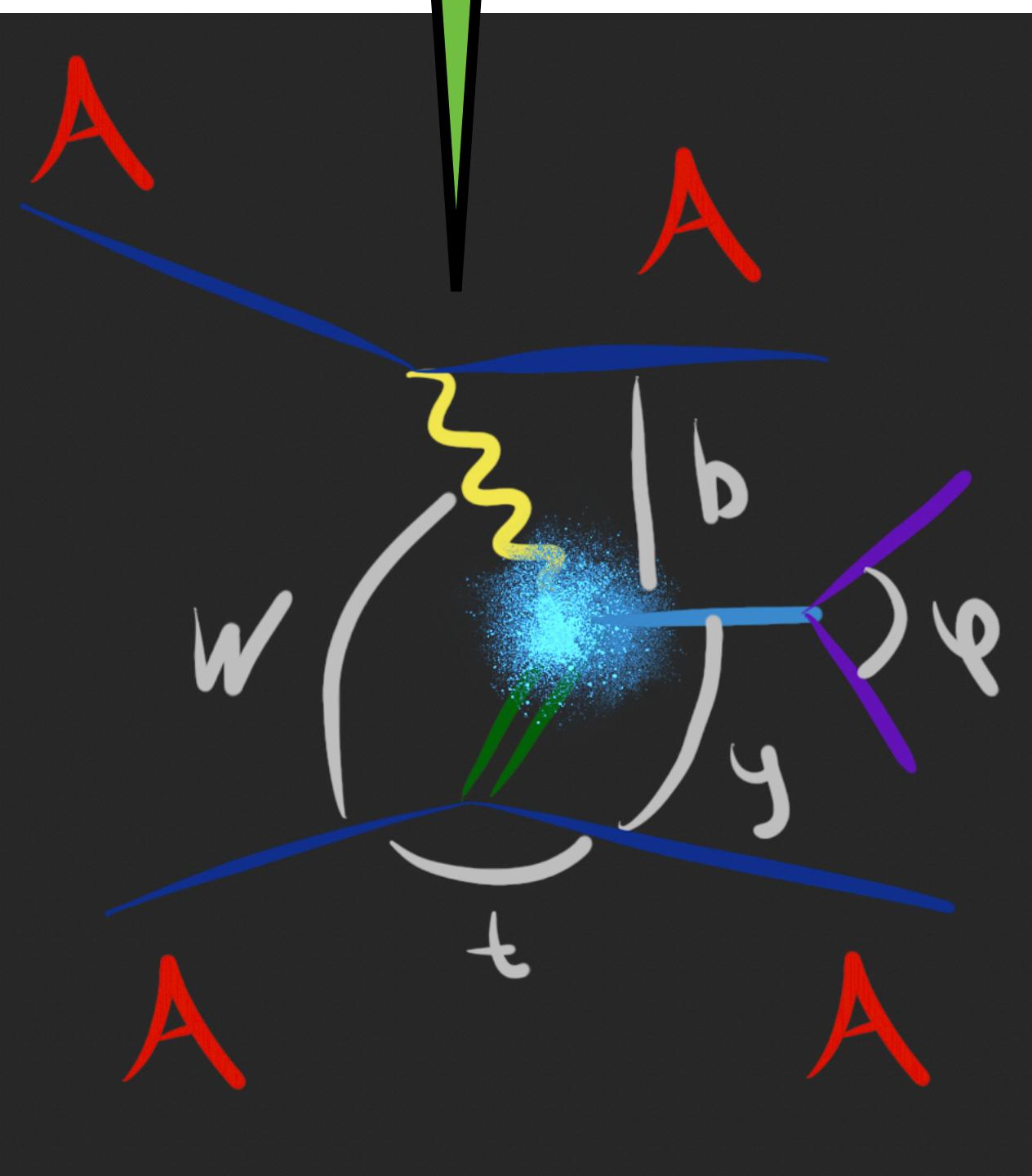
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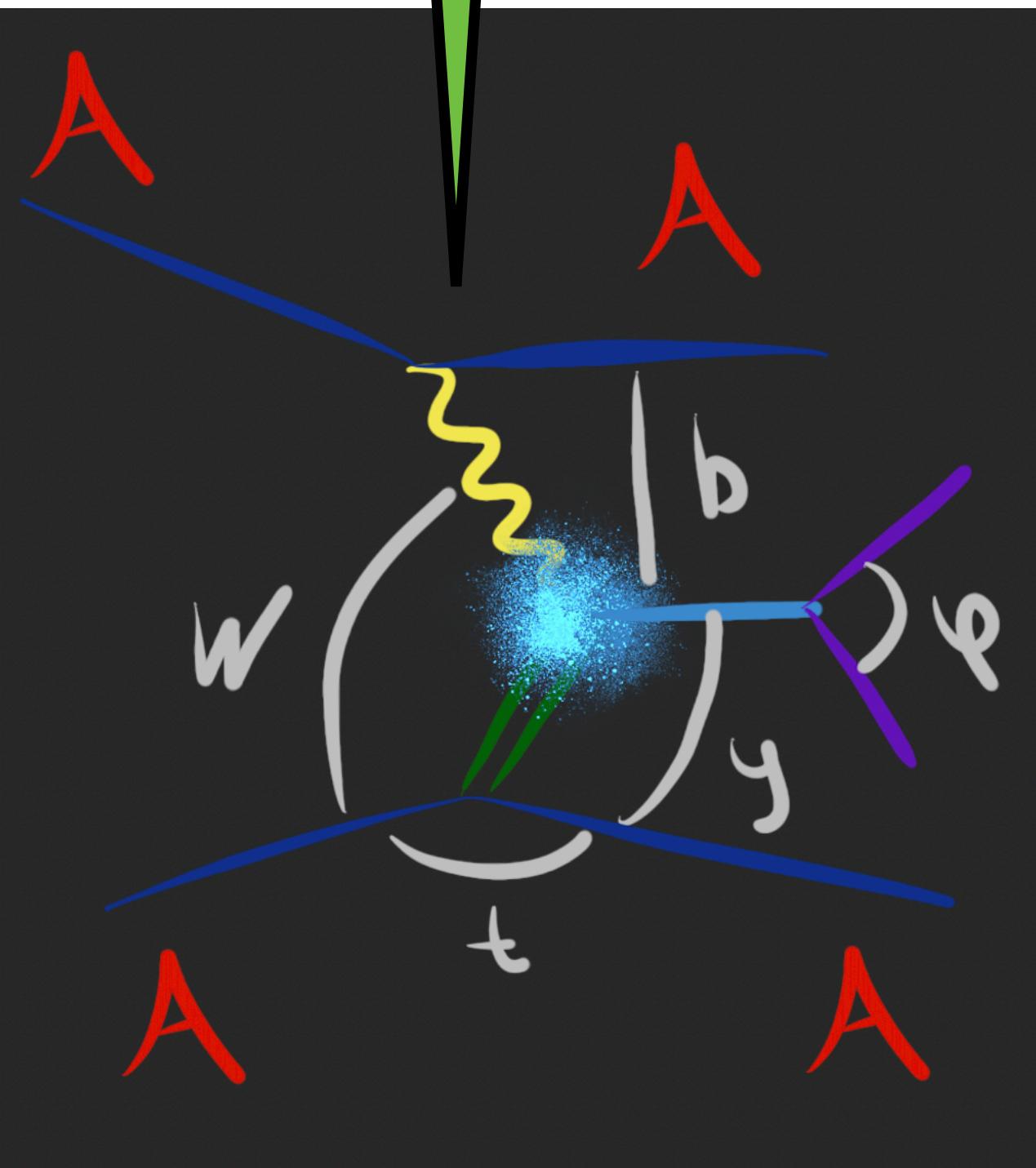
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Linear polarisation along the impact parameter

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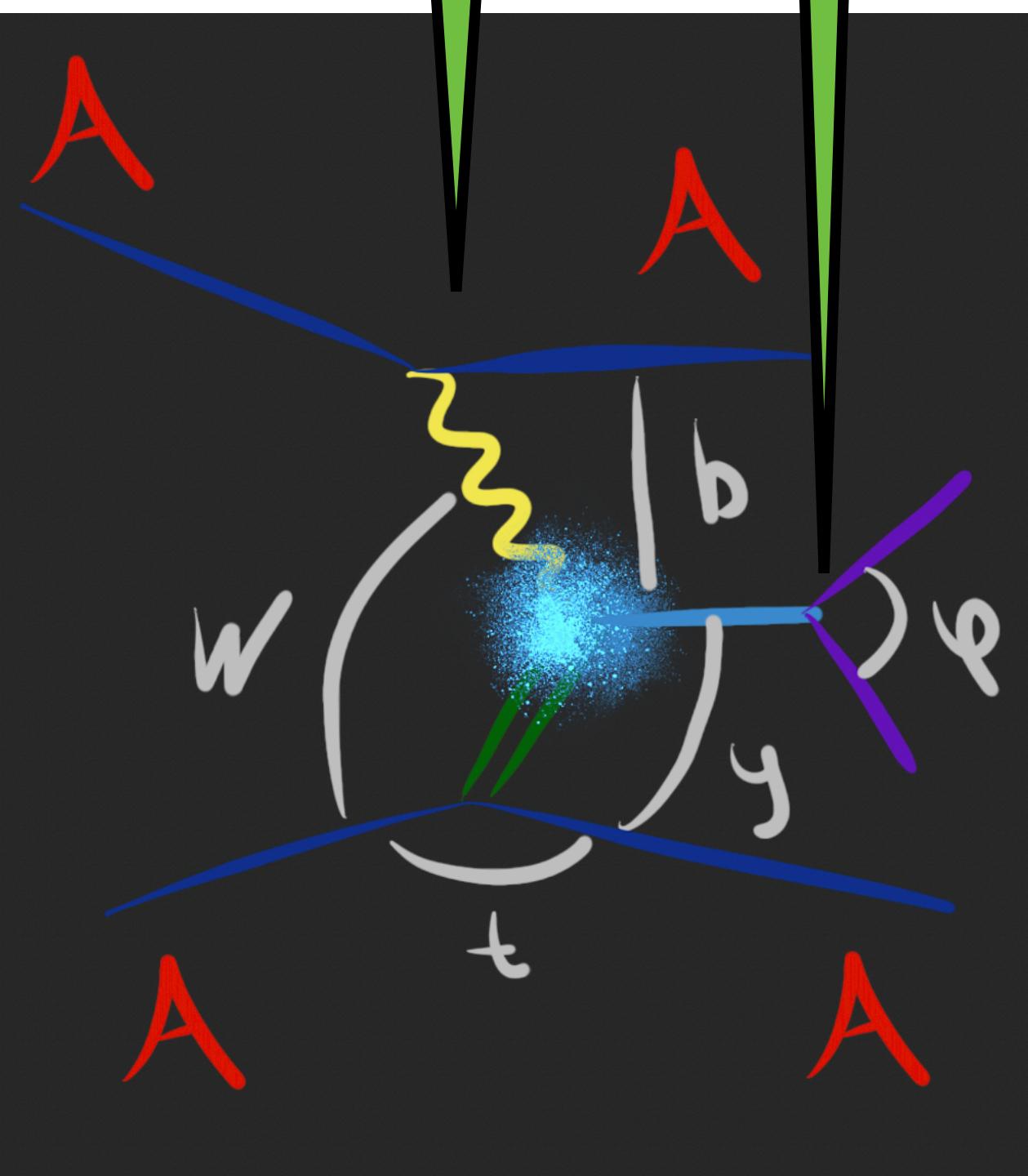


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In photo-production the polarisation is transferred to the **vector meson**

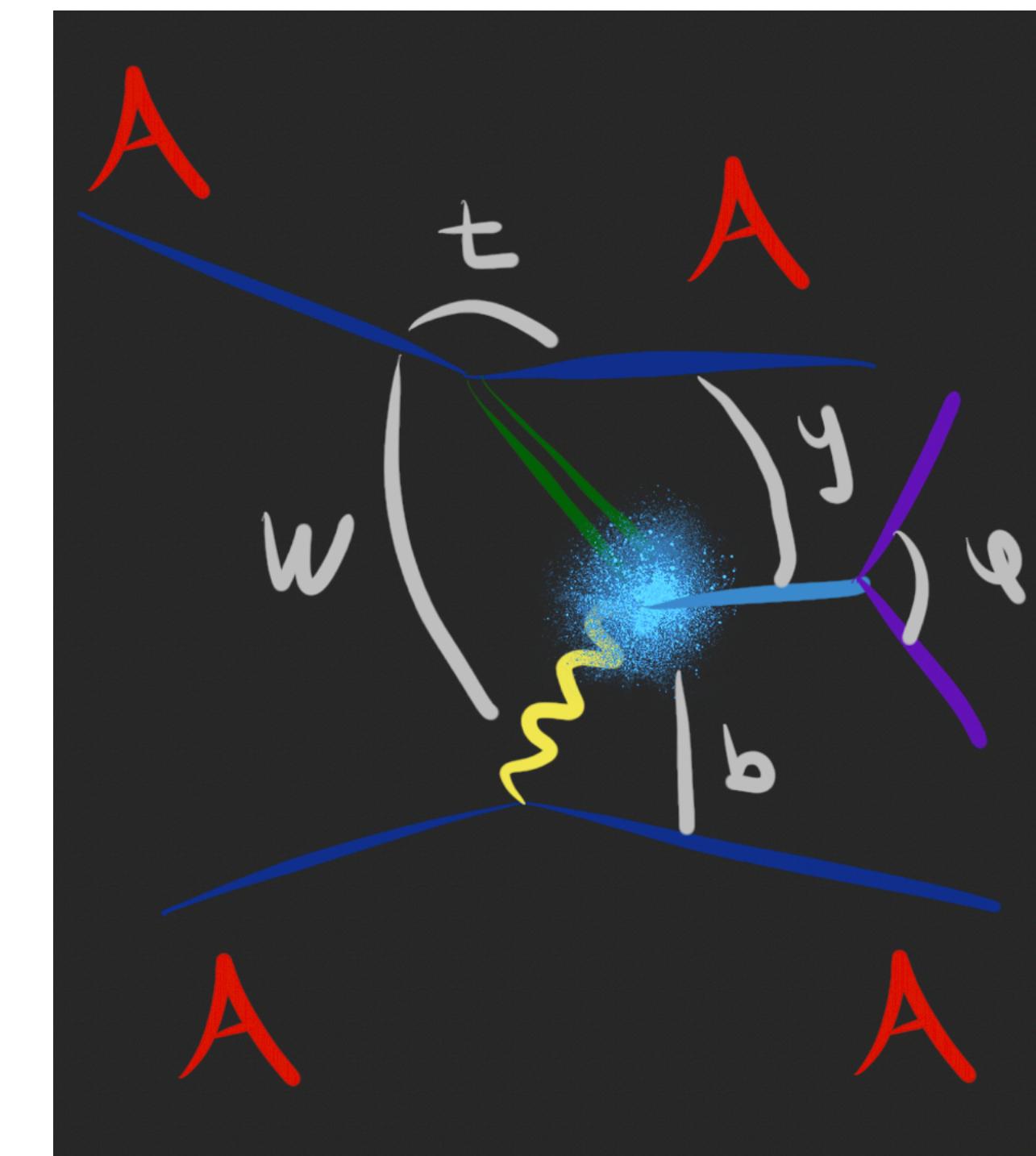
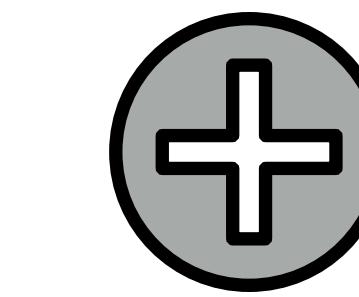
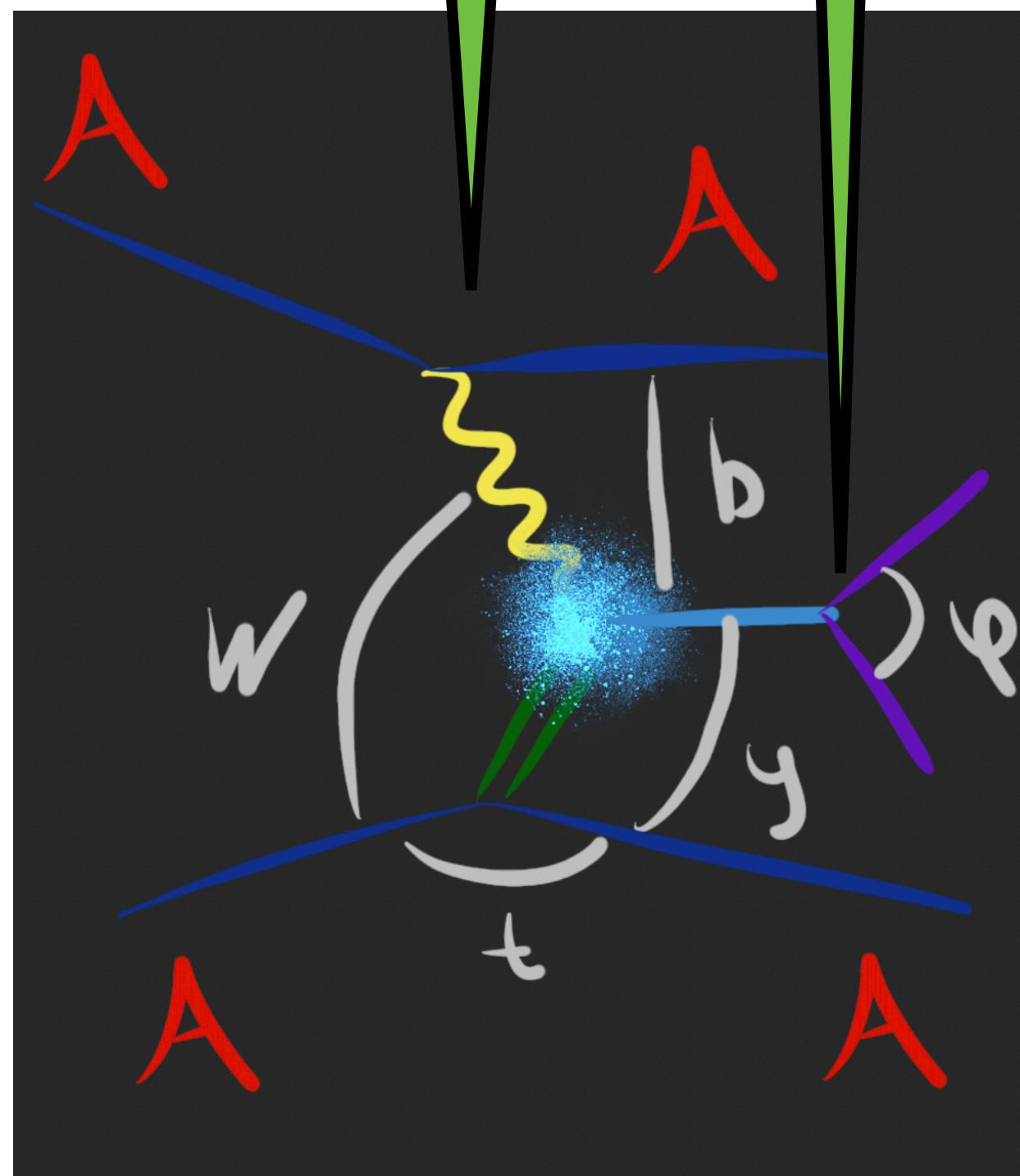


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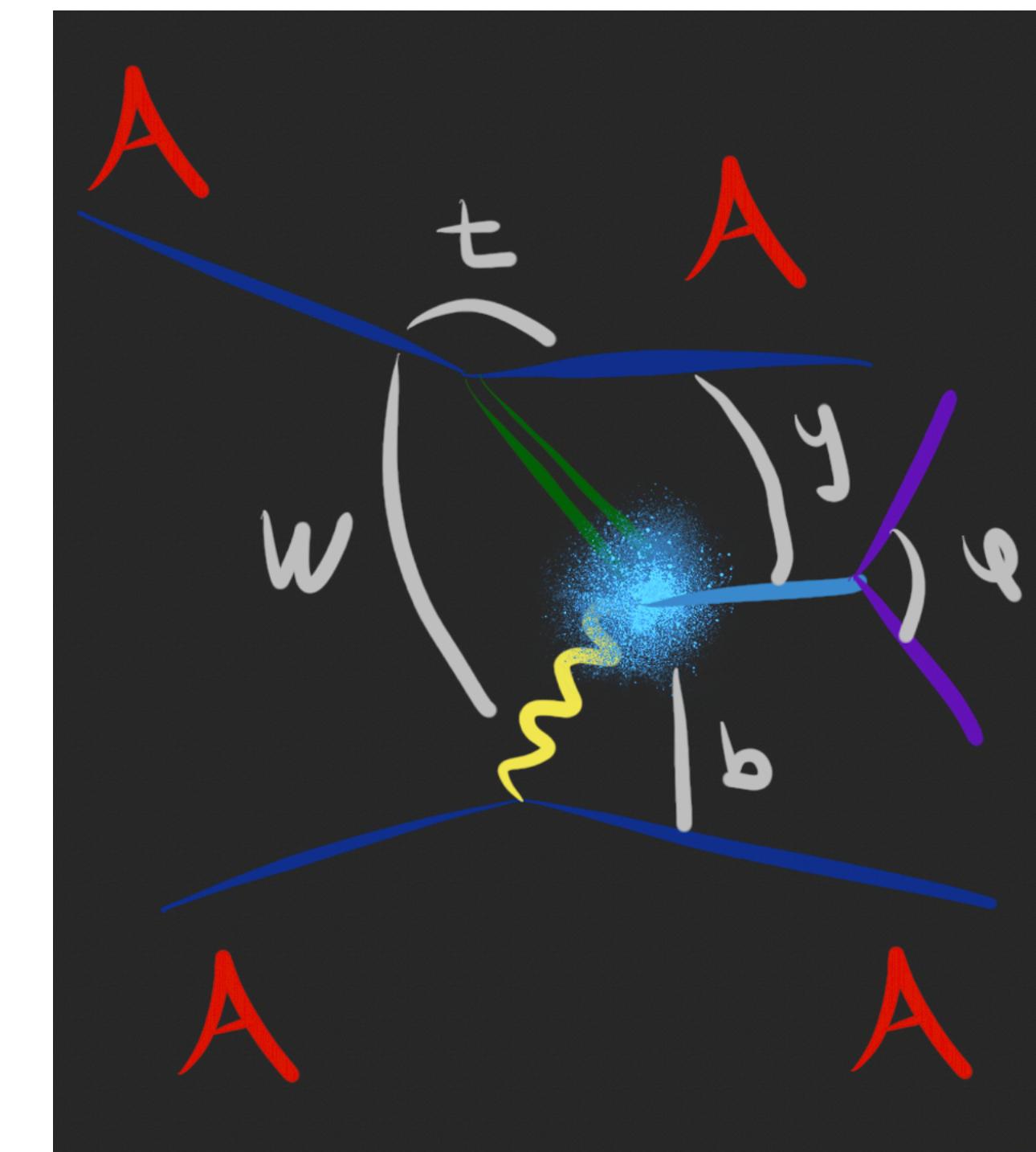
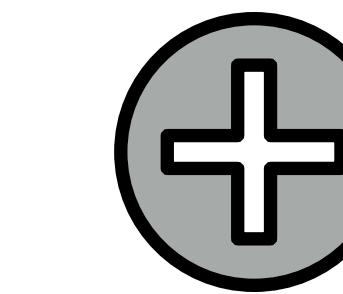
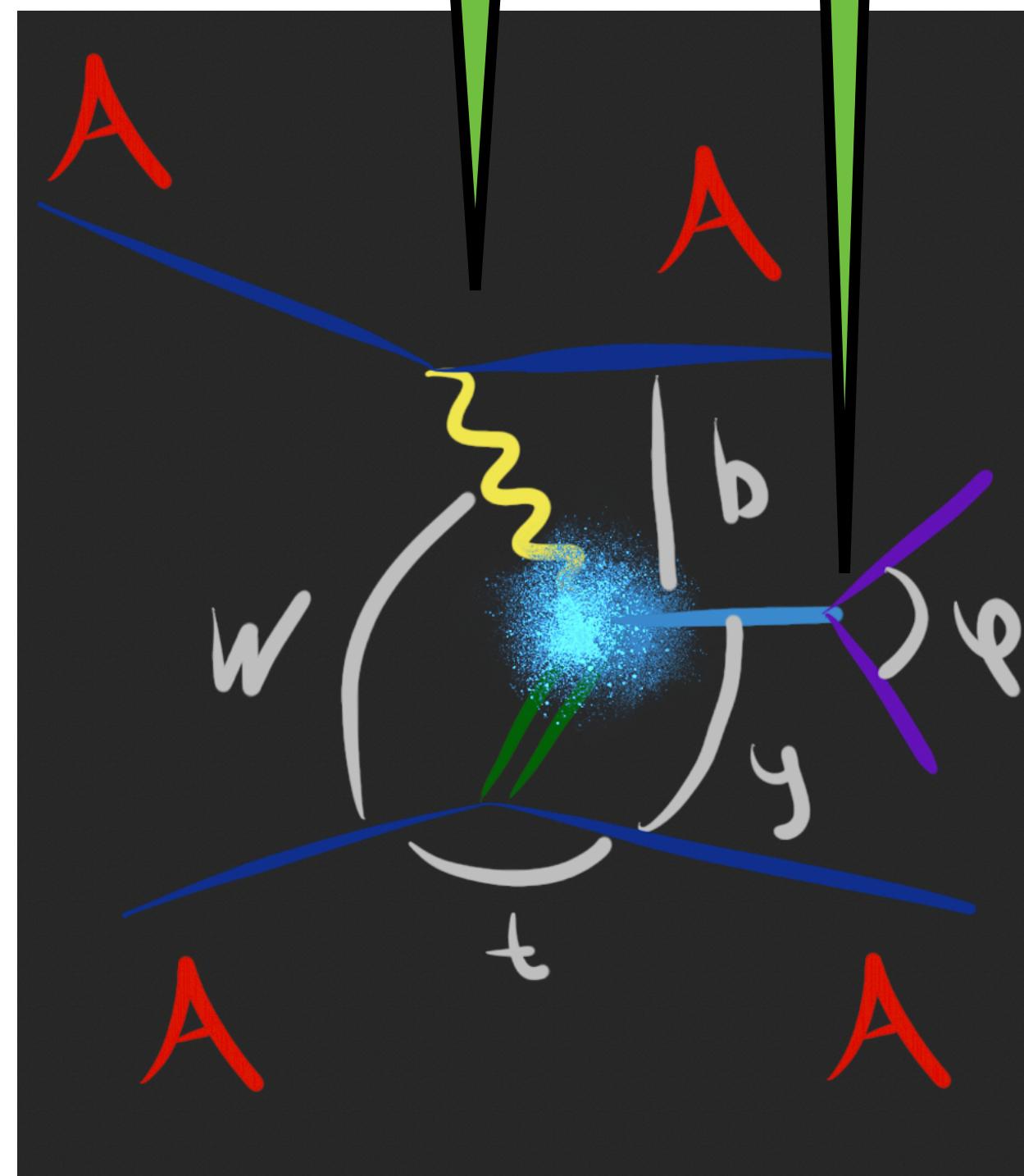
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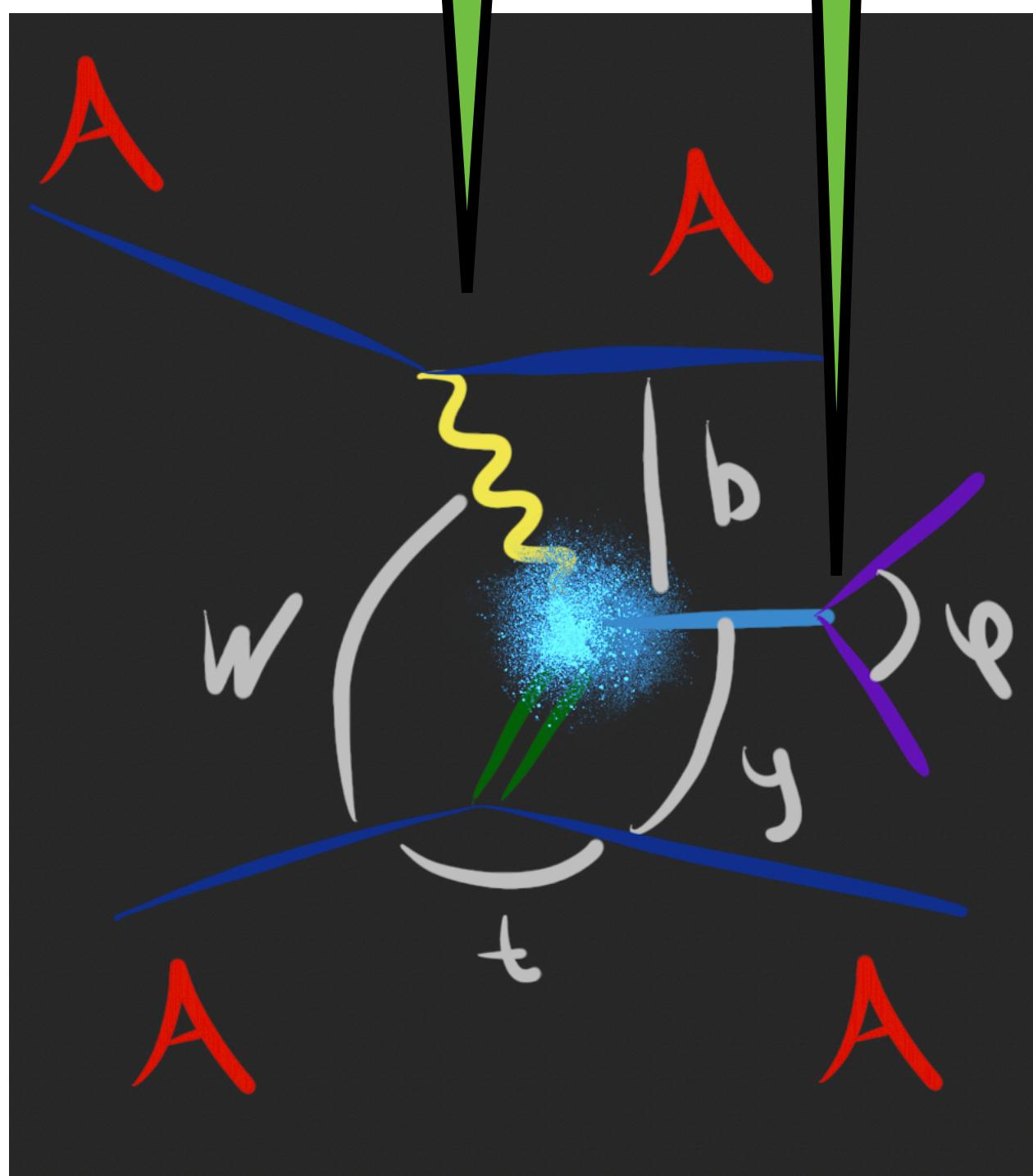
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$\phi$  is the **azimuth** angle between the sum and the difference of the **pion** four-vectors

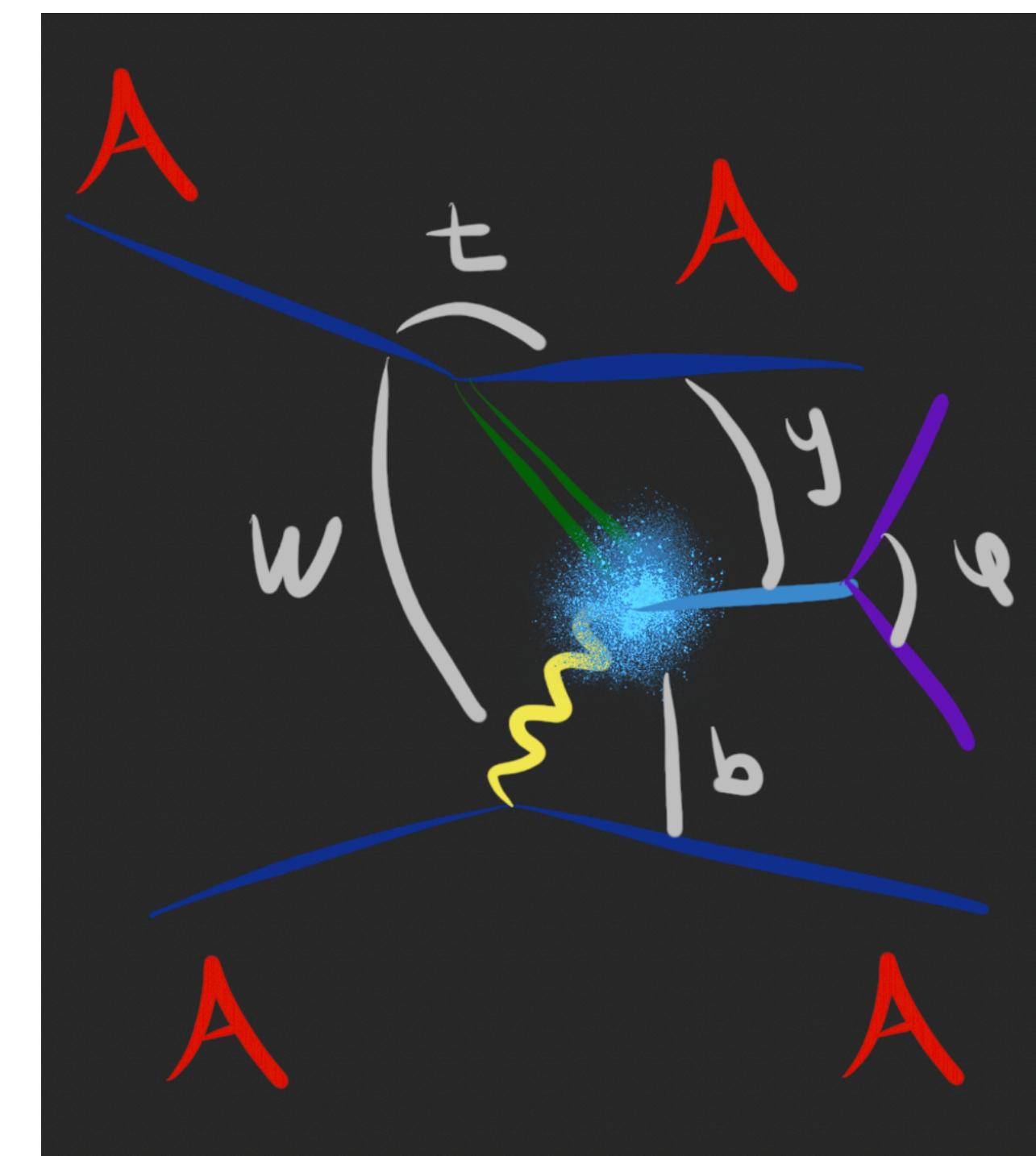
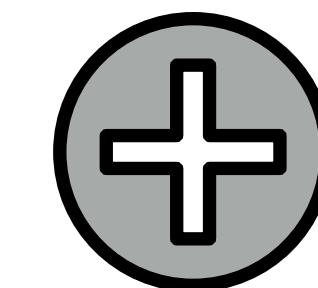
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Linear polarisation along the impact parameter

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Interference effects are more **important** at mid-rapidity, where both amplitudes are similar, and at small impact parameters

**Interference** and the decay into scalar particles produces a  $\cos(2\phi)$  modulation

$\phi$  is the **azimuth** angle between the sum and the difference of the **pion** four-vectors

# Data analysis

## Data

2015 Pb-Pb data at 5.02 TeV  
CUP9 trigger (vetoes in AD/V0 and 0STP)  
Lumi  $\approx 0.48 \text{ 1}/\mu\text{b}$

Paper: <https://alice-publications.web.cern.ch/node/9523>

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Standard event, track and pion-pair selection  
Three neutron classes: 0n0n, 0nXn, XnXn  
15  $\phi$  bins in each class  
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Soeding-like invariant mass fit in each bin  
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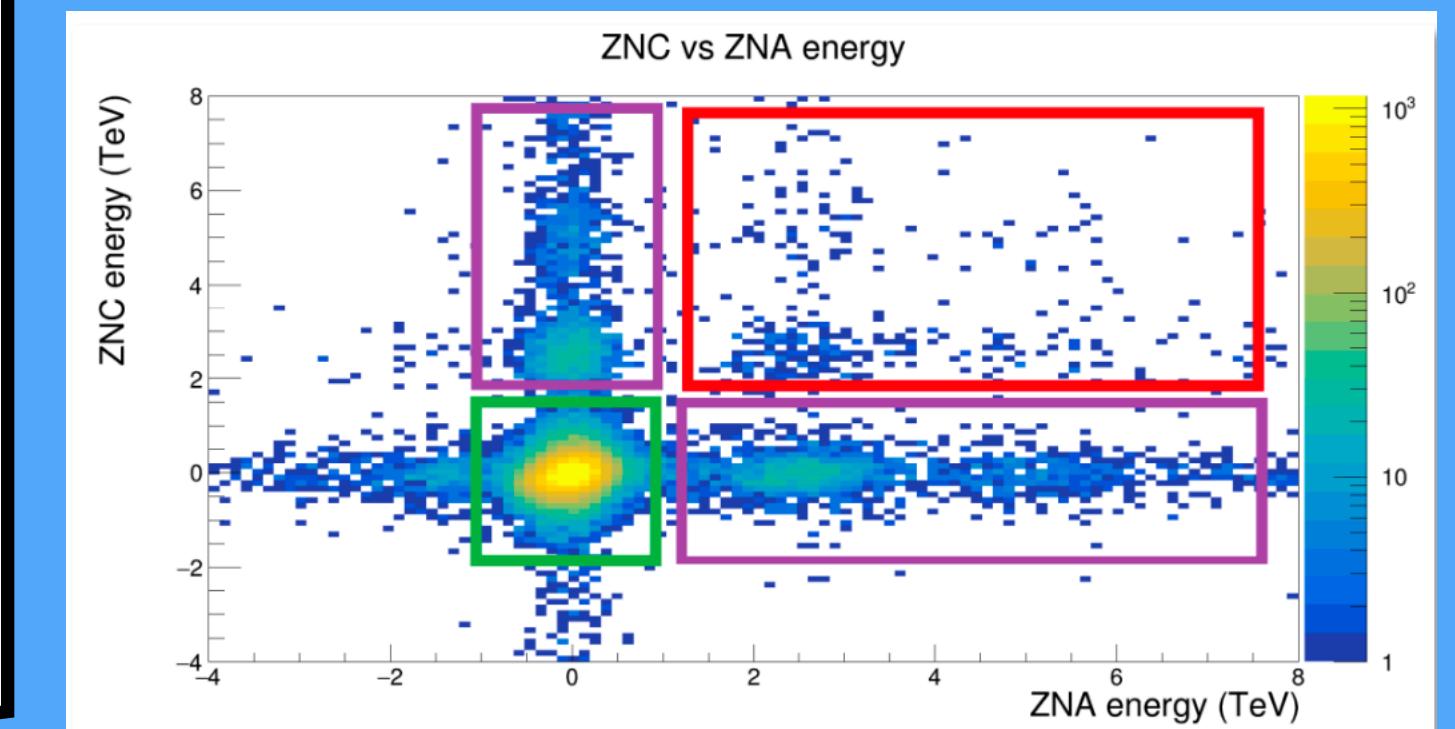
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## Classes defined with ZN



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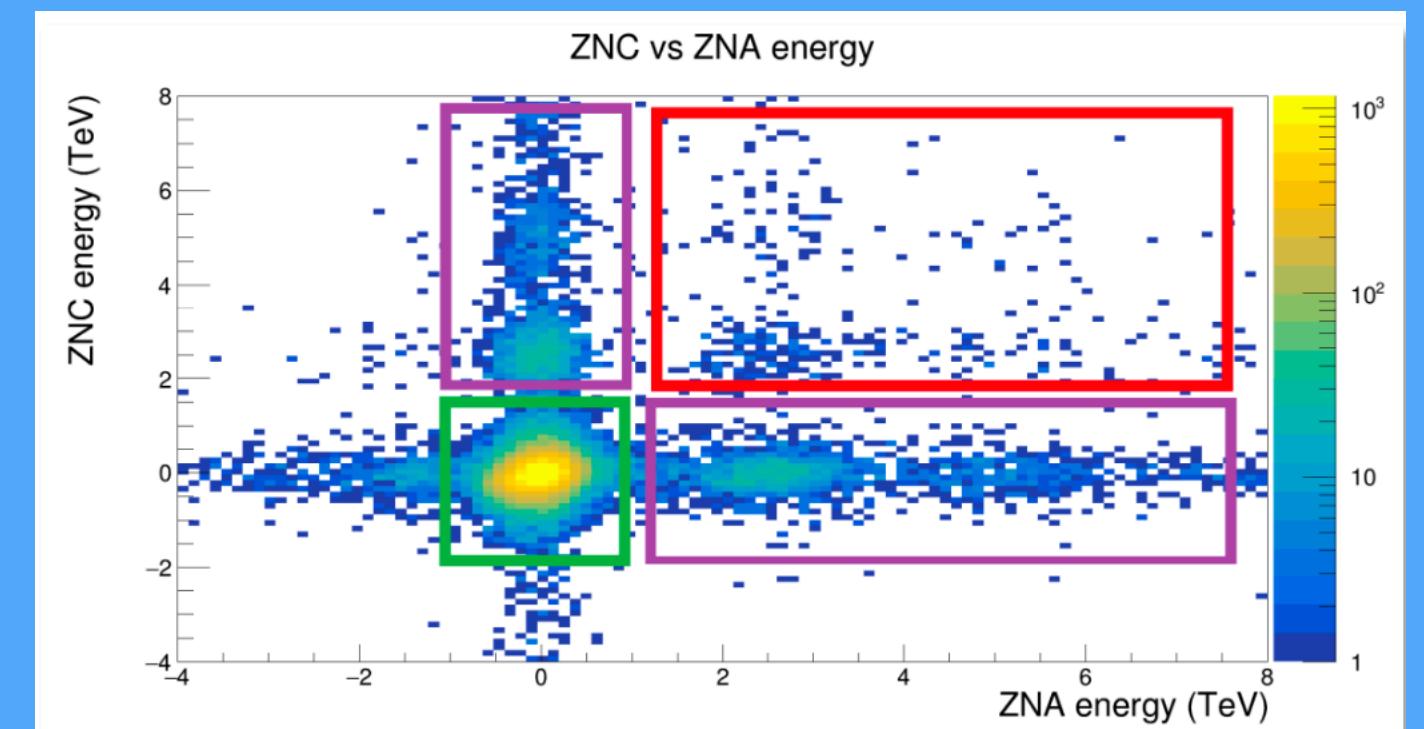
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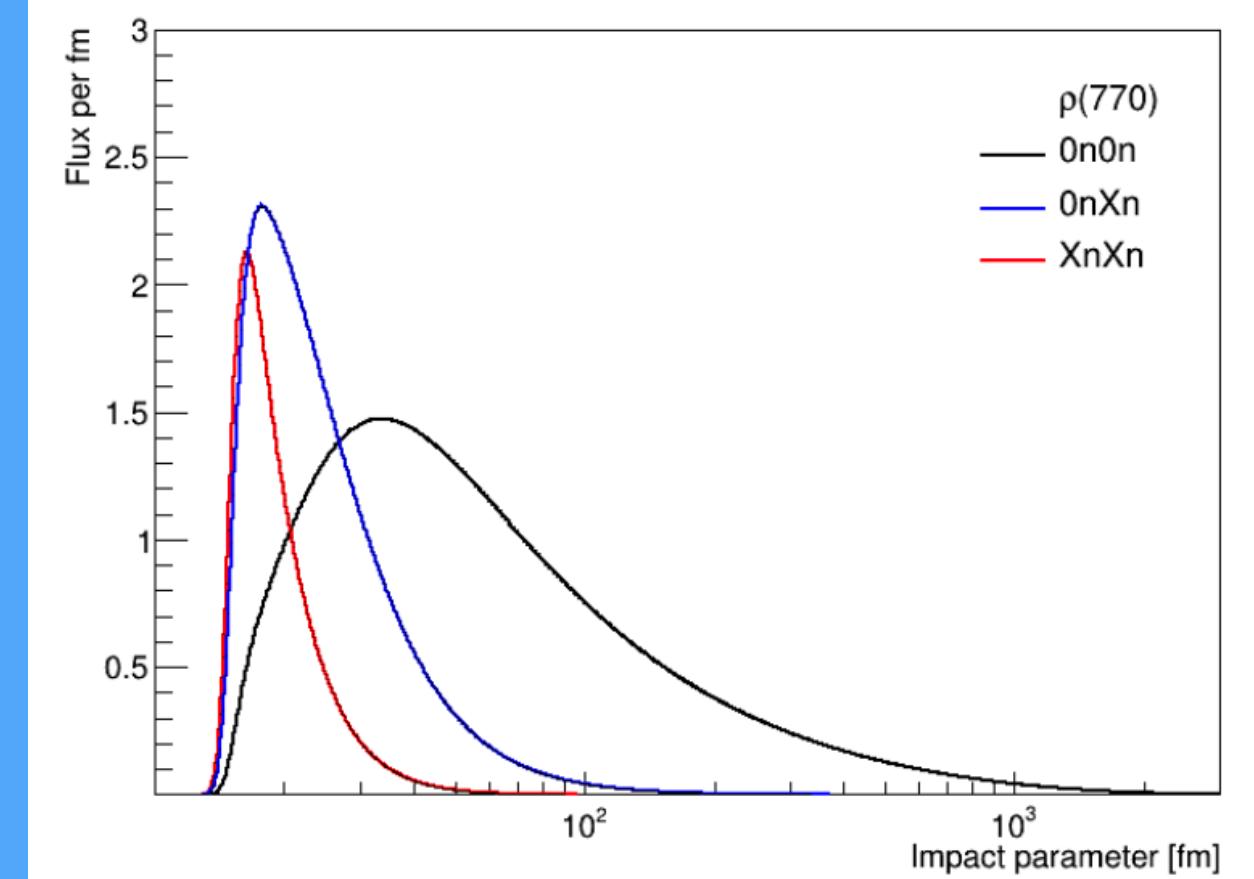
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## Selection of impact-parameter ranges



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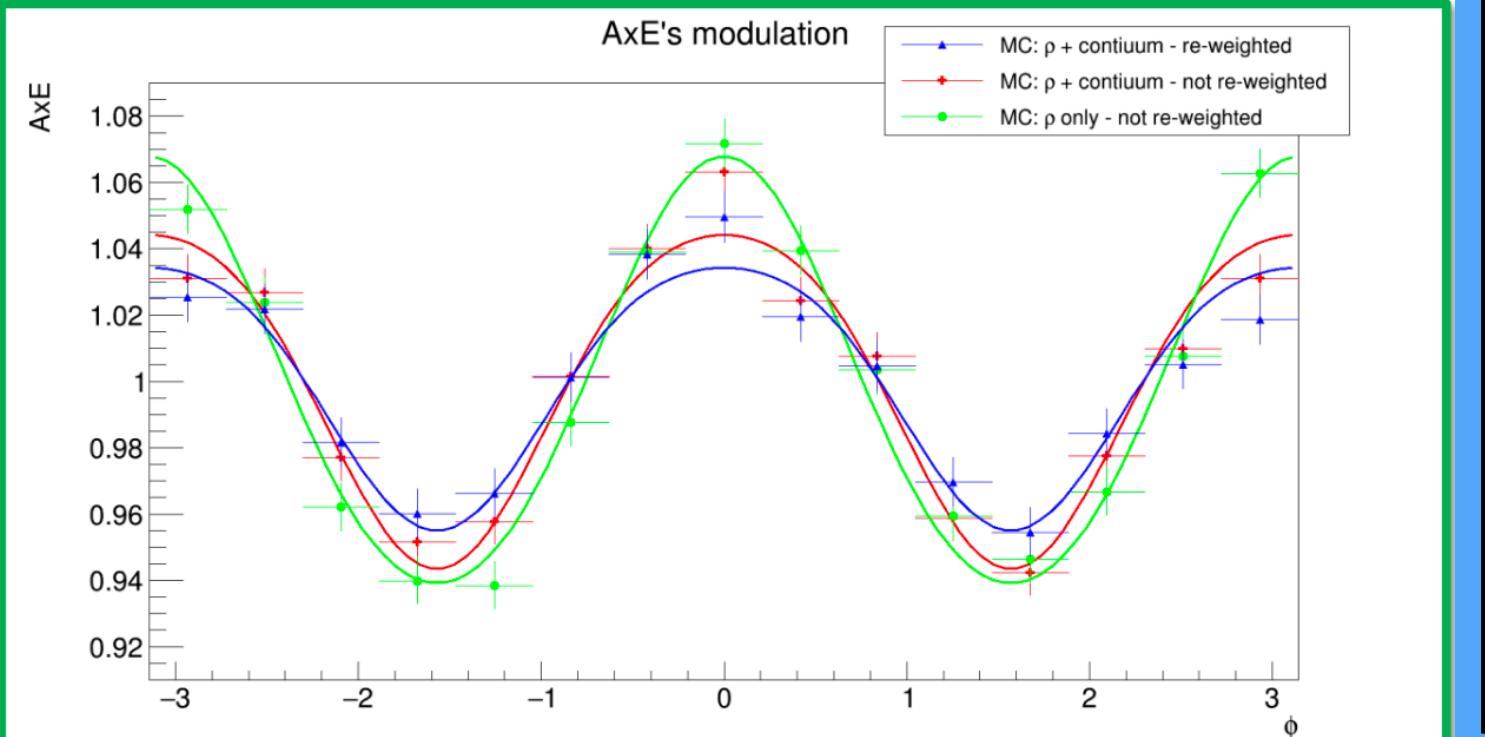
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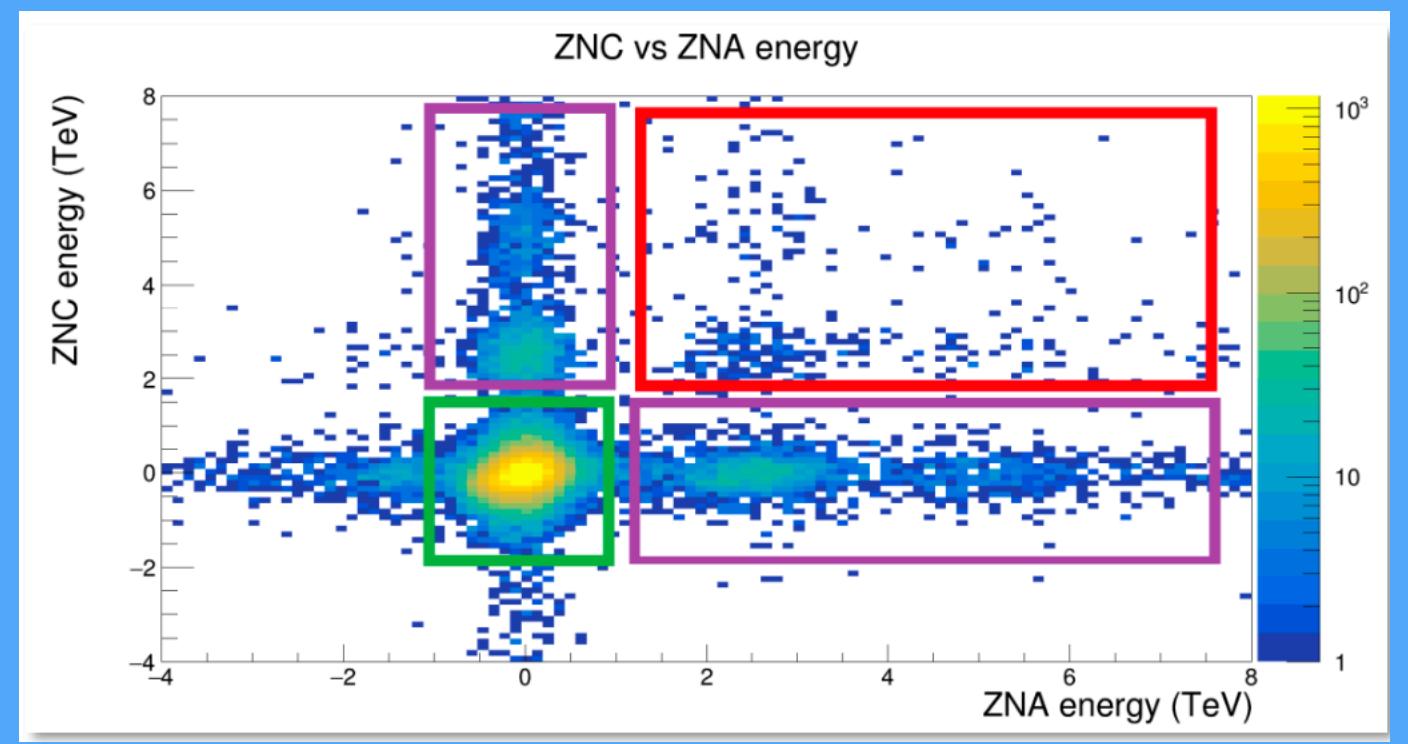
STARlight needed reweighting



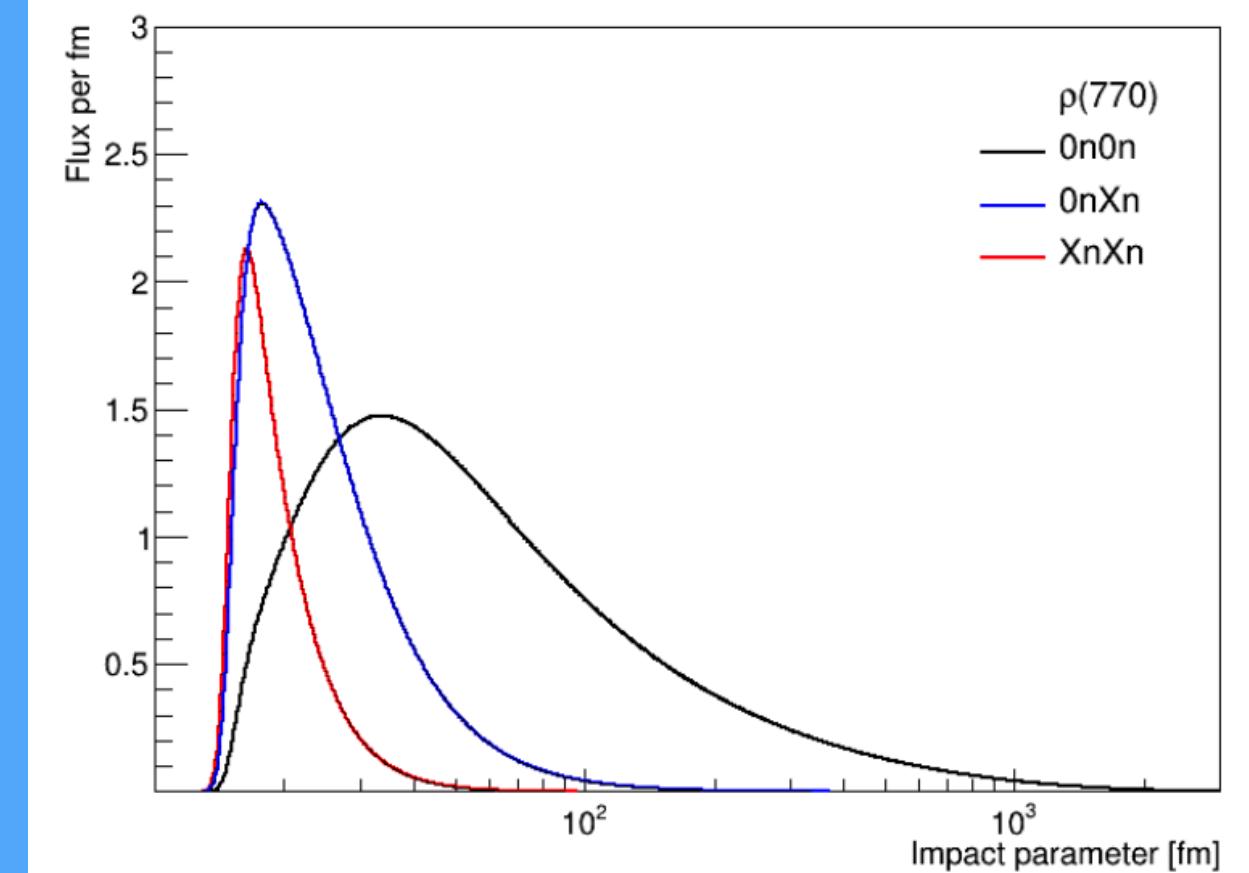
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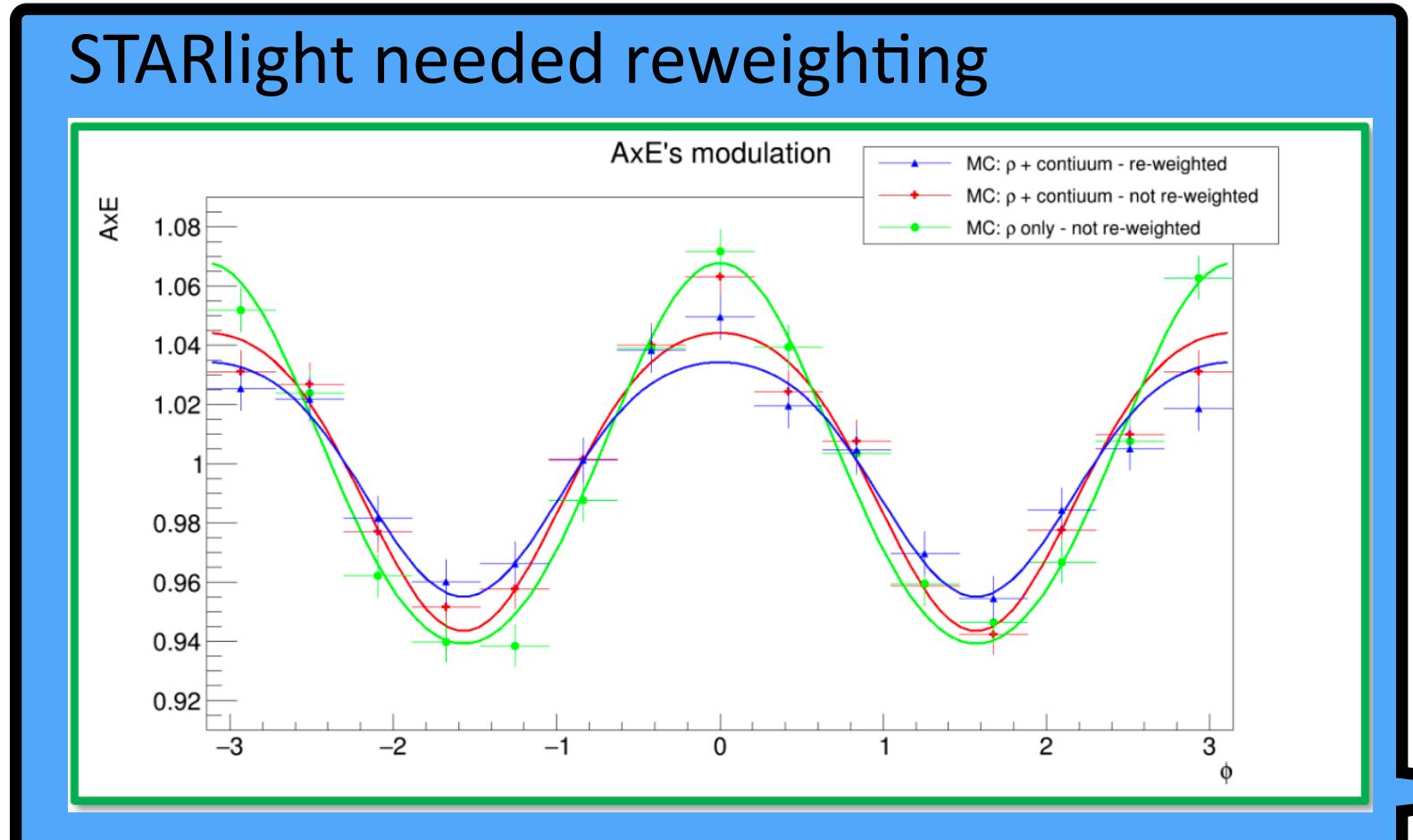
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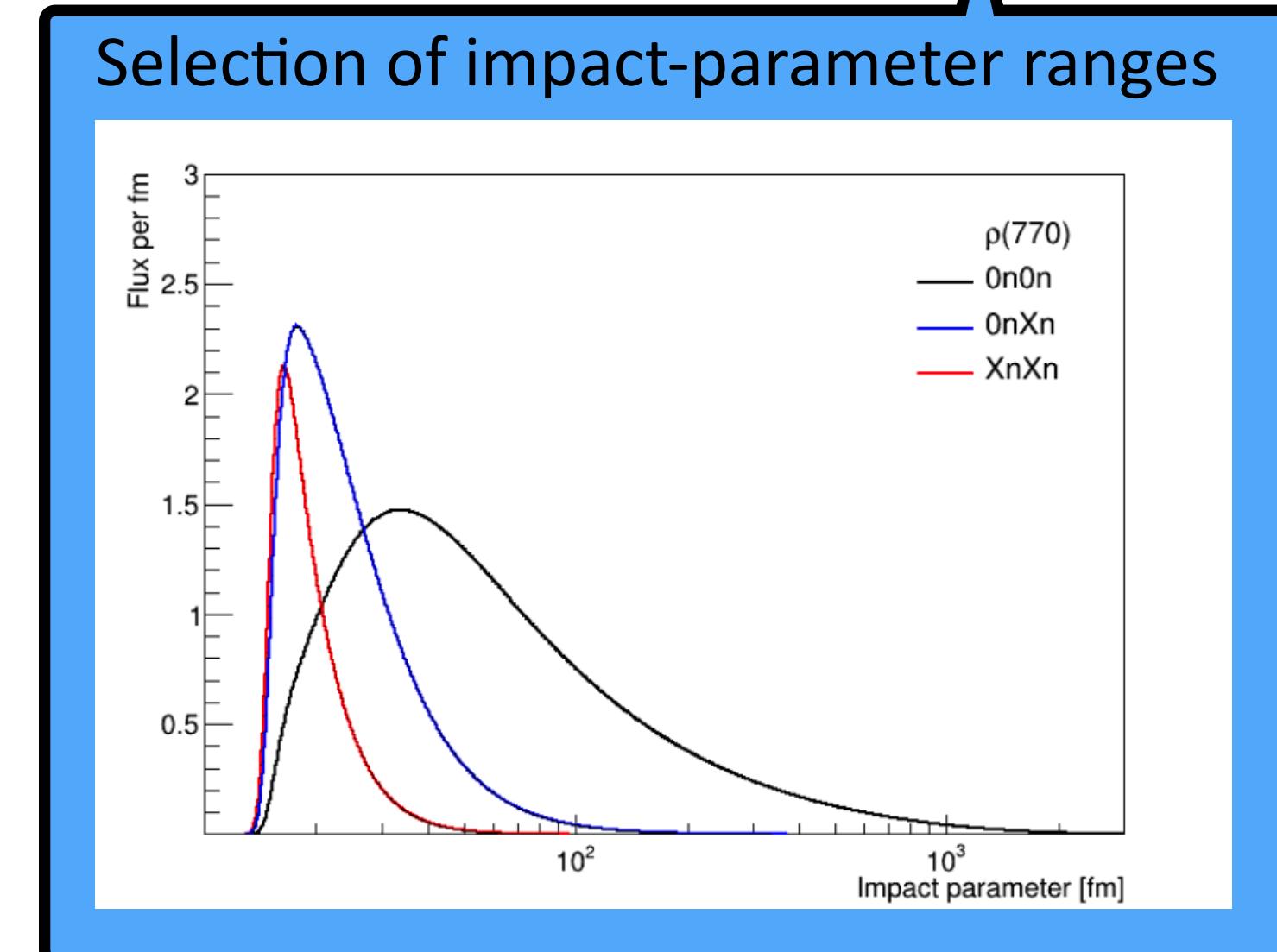
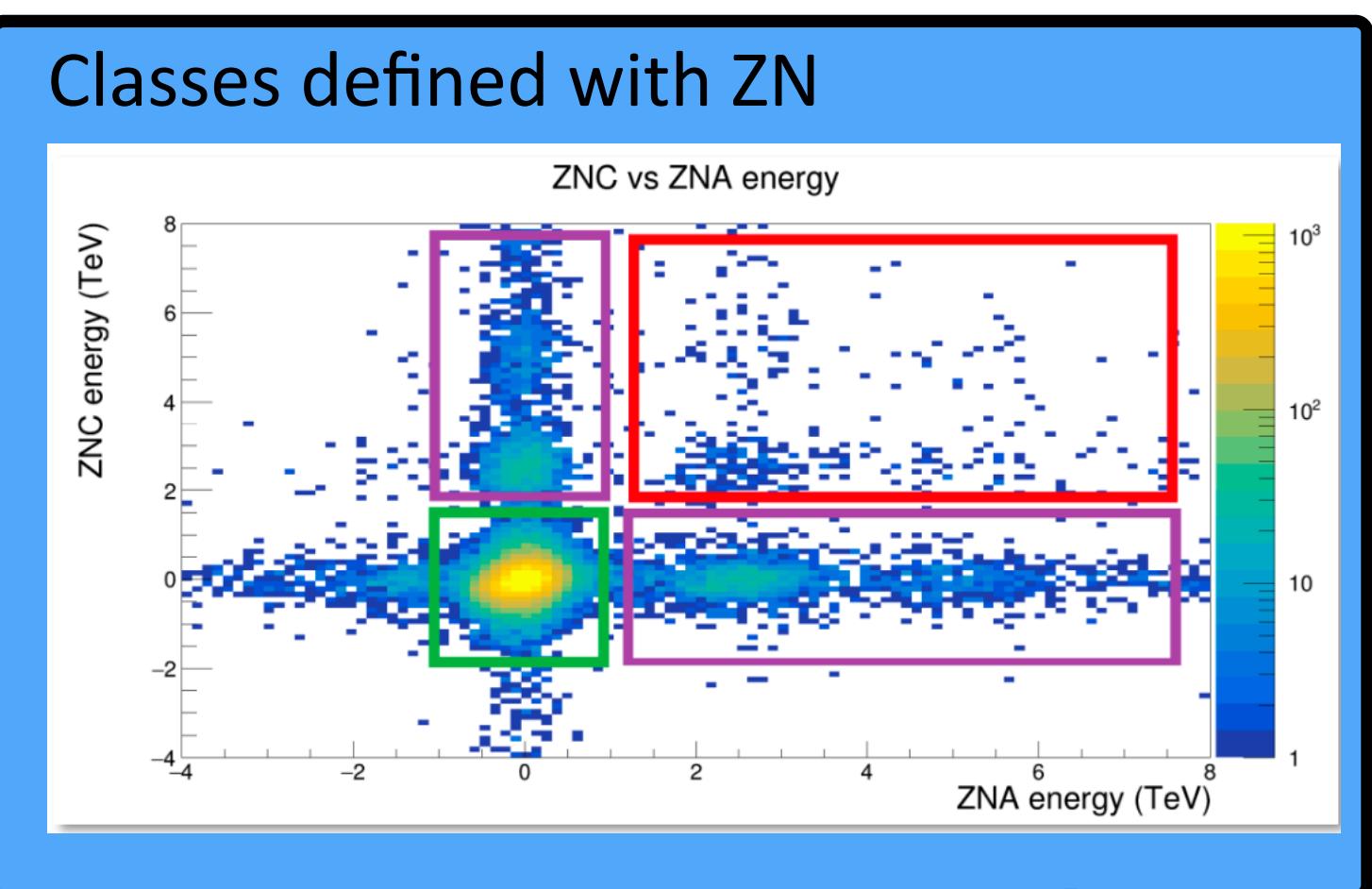
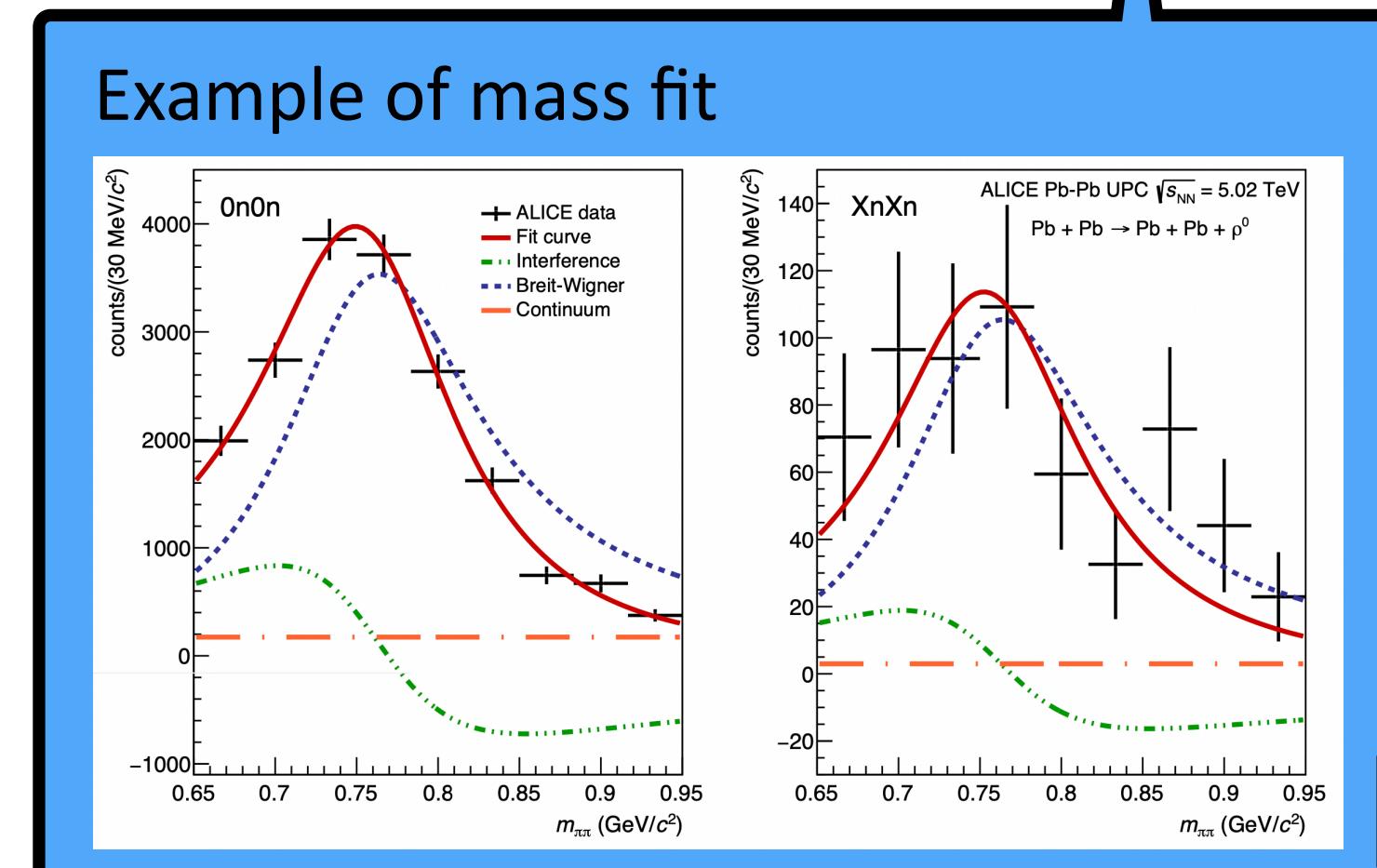
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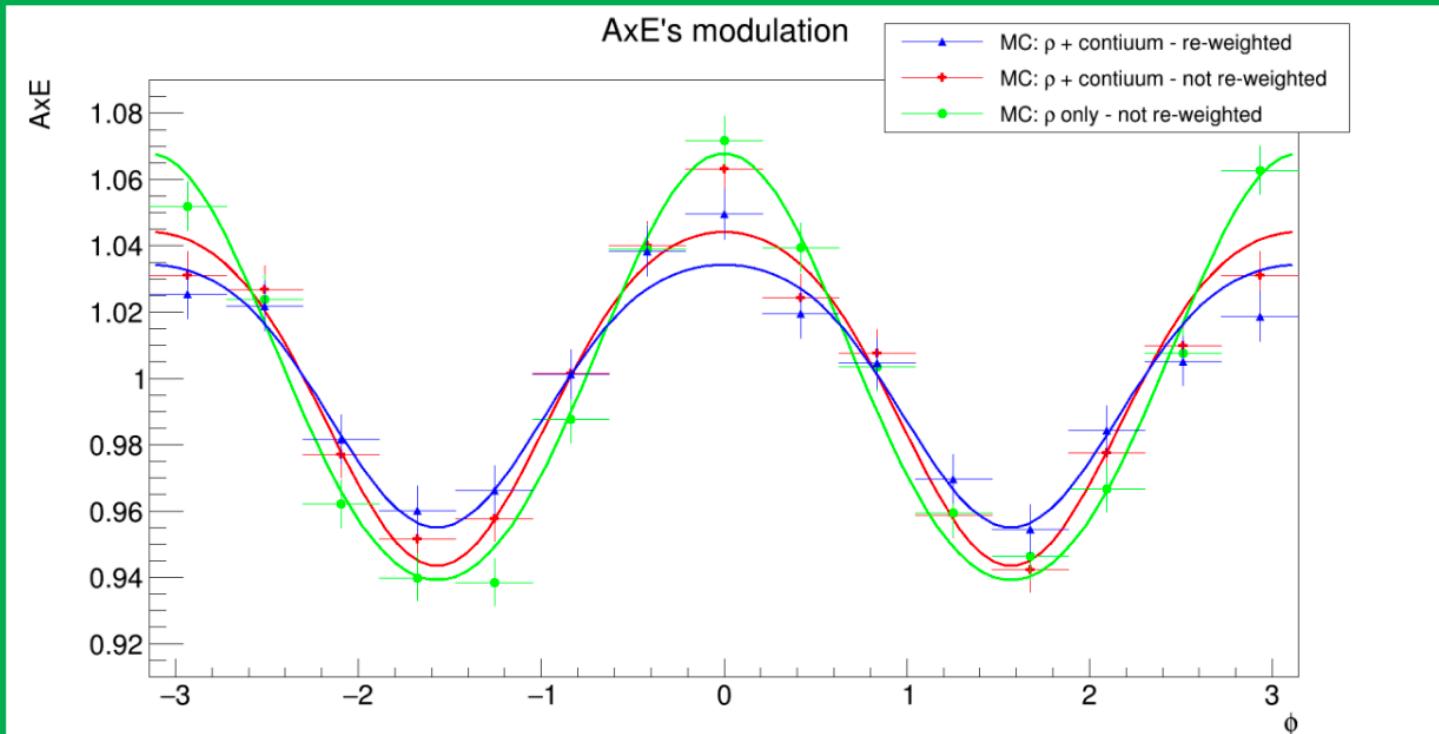
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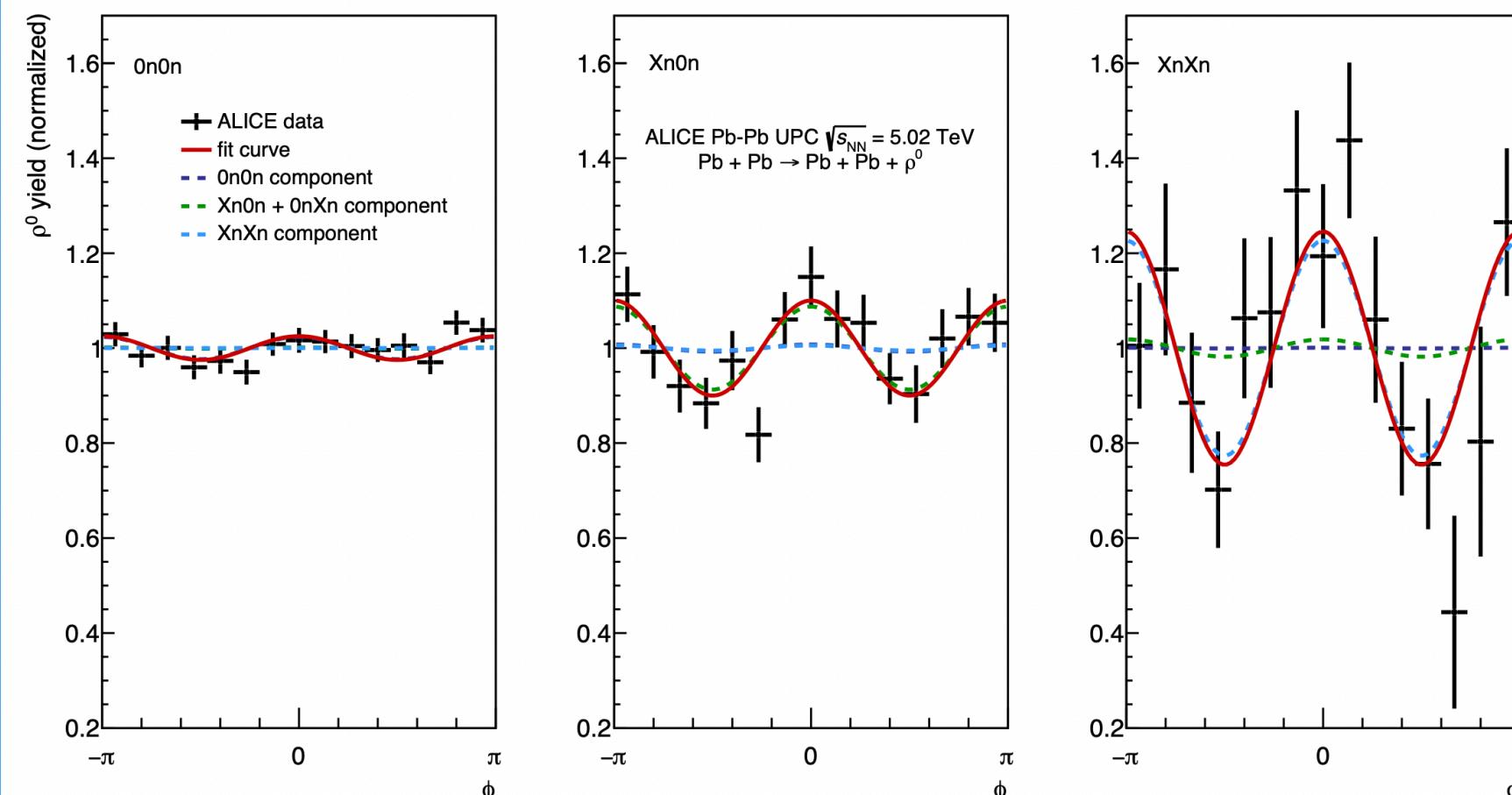
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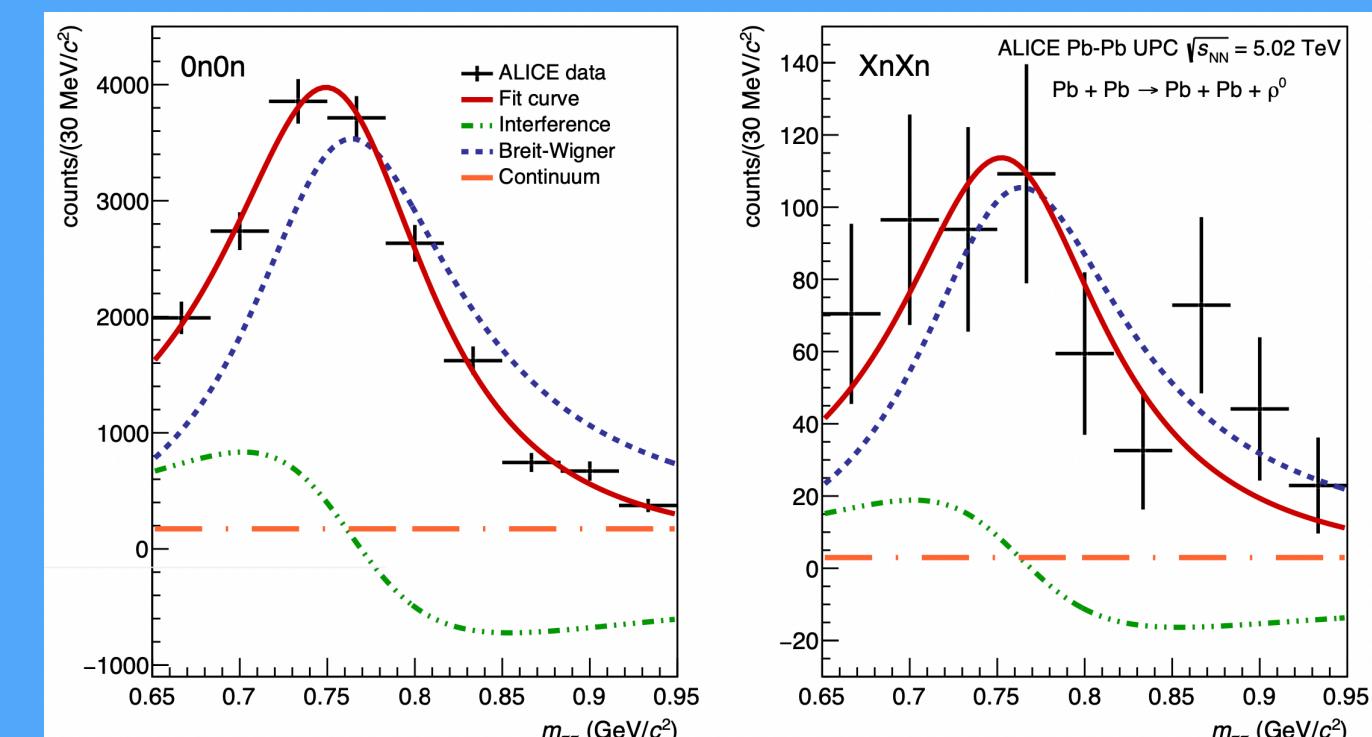
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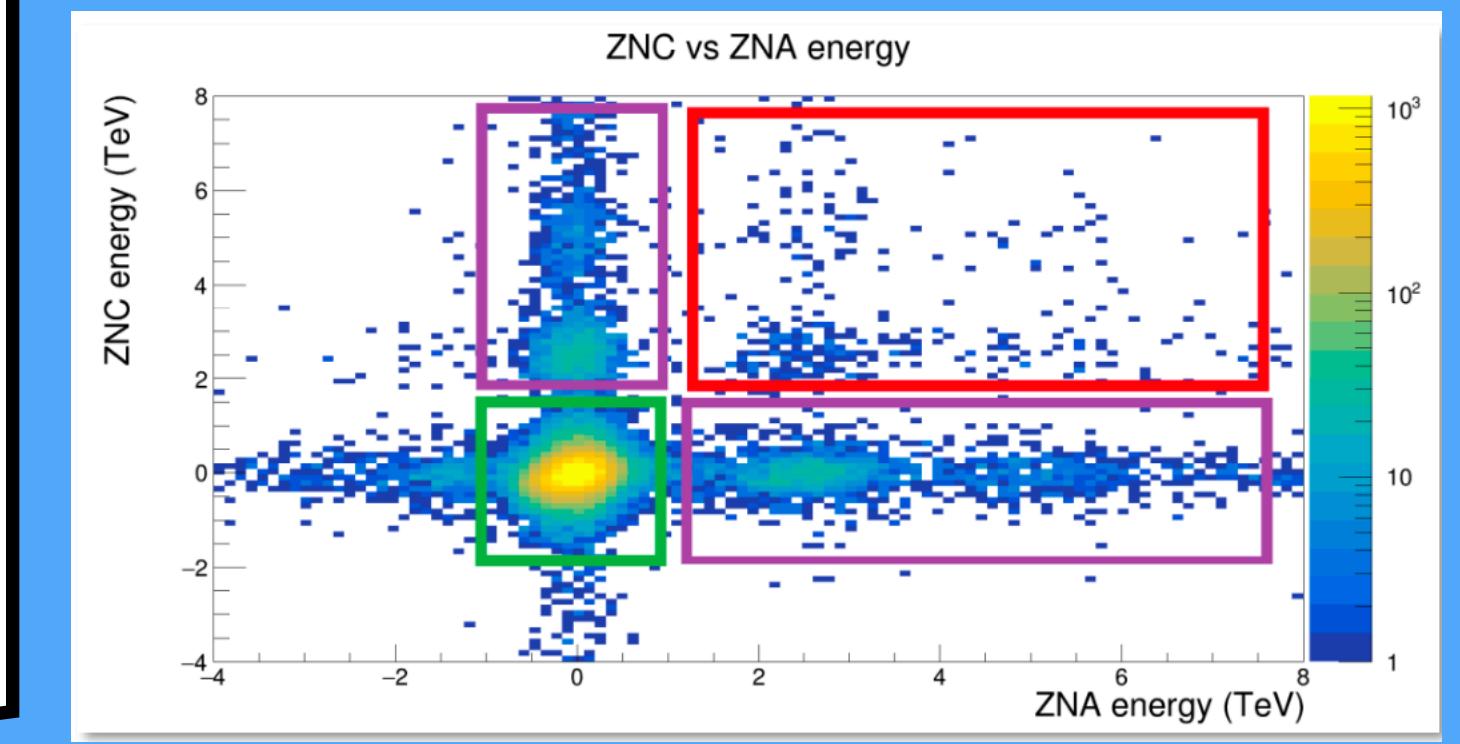
## Simultaneous fit to account for migrations



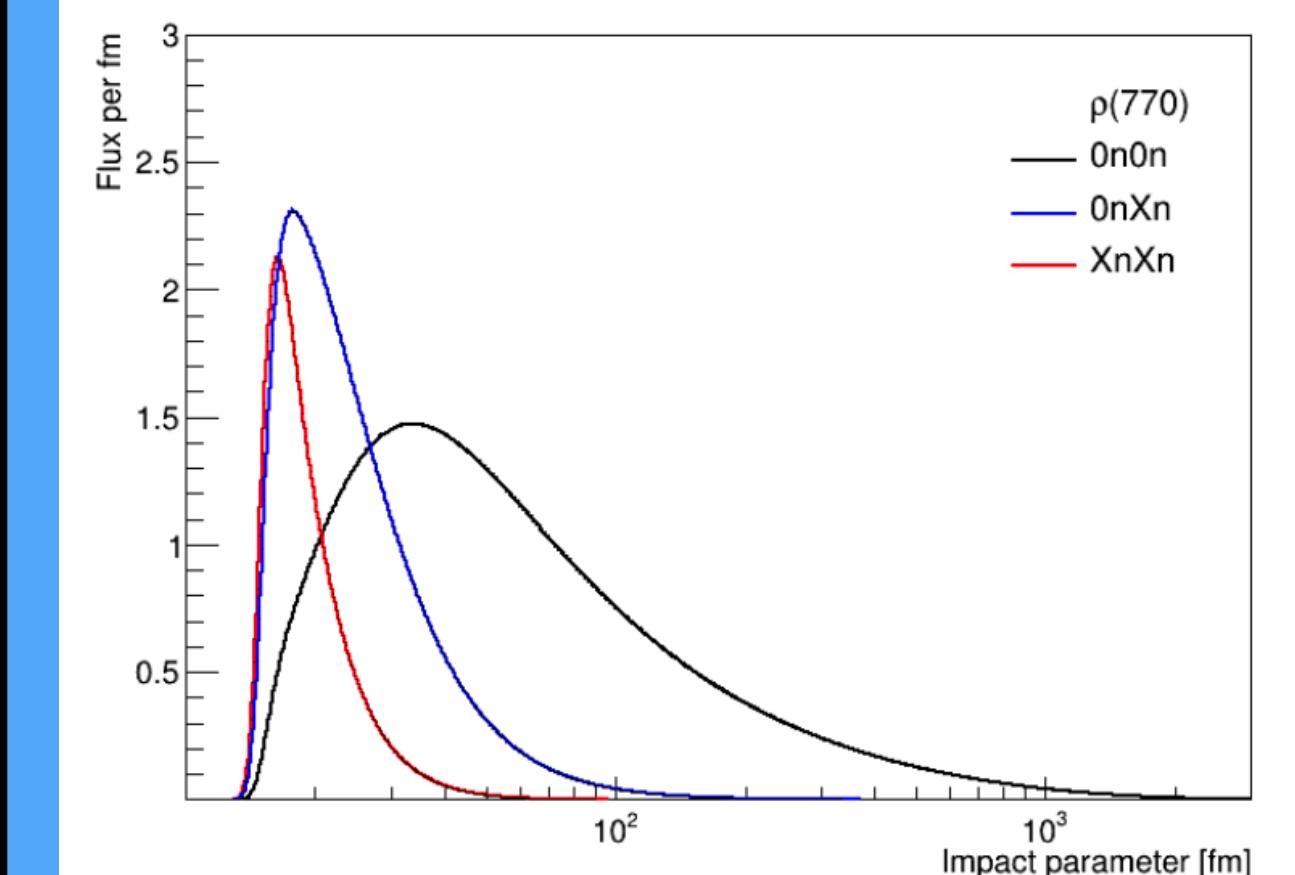
## Example of mass fit



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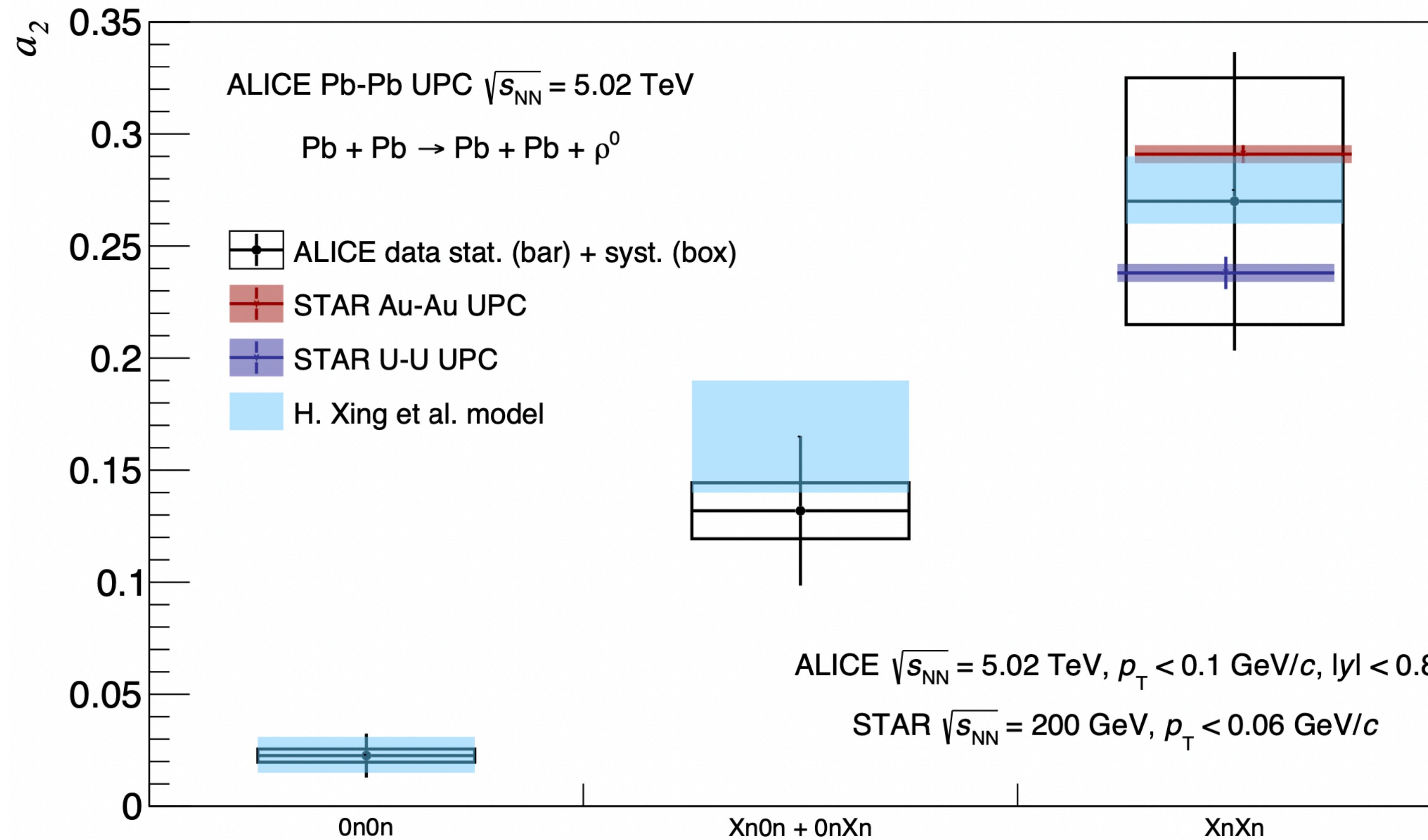


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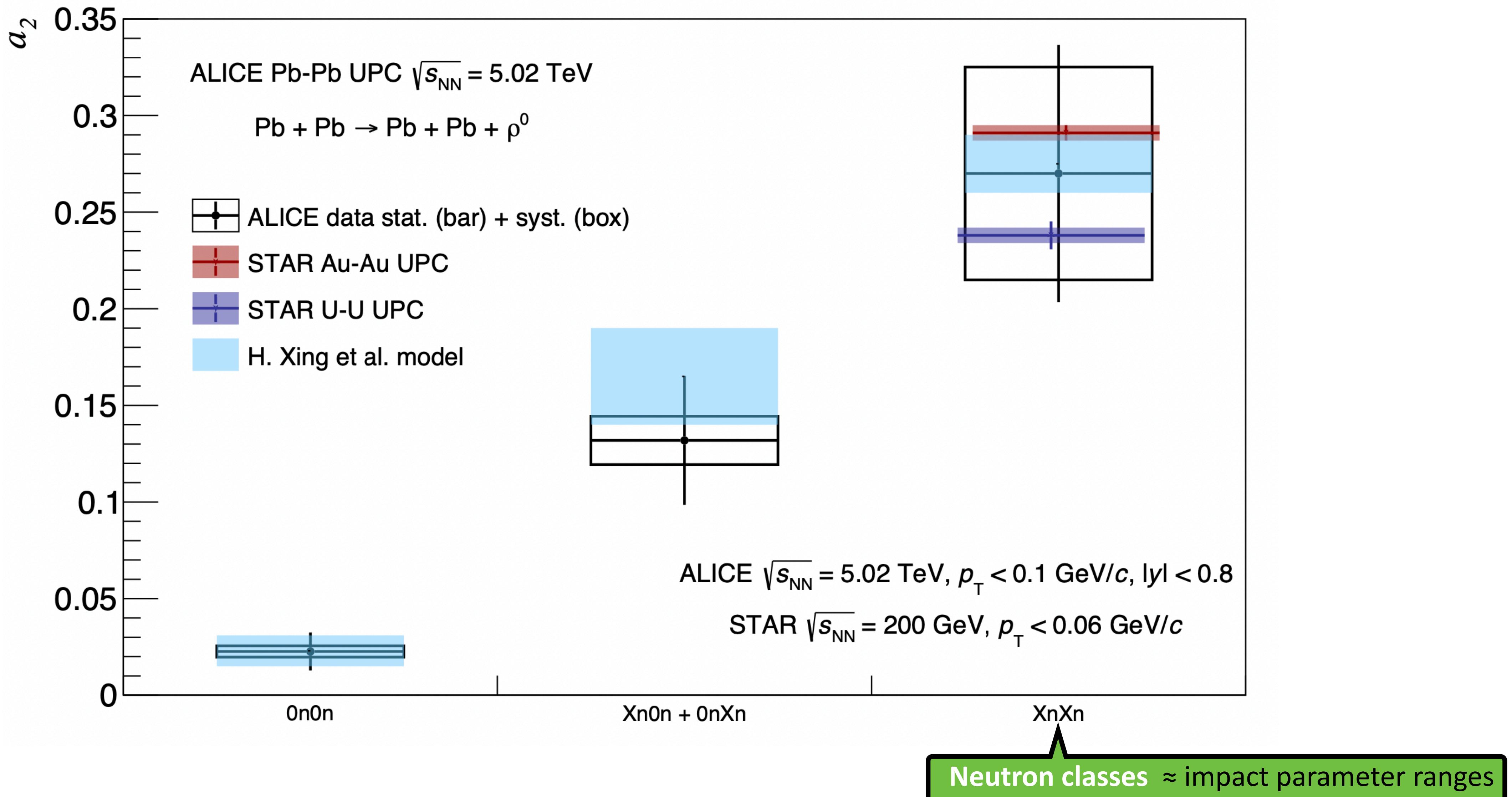


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# Results

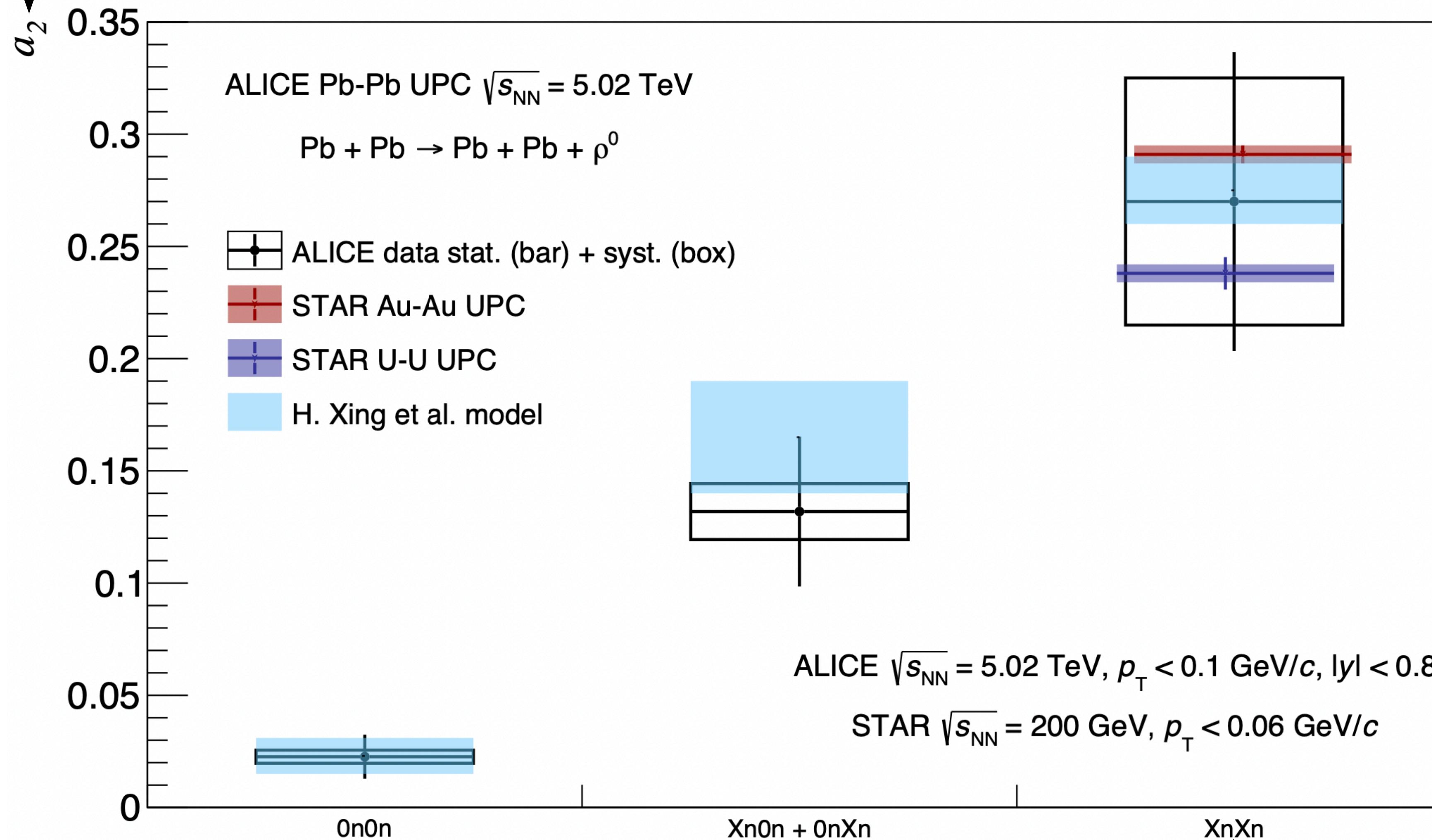


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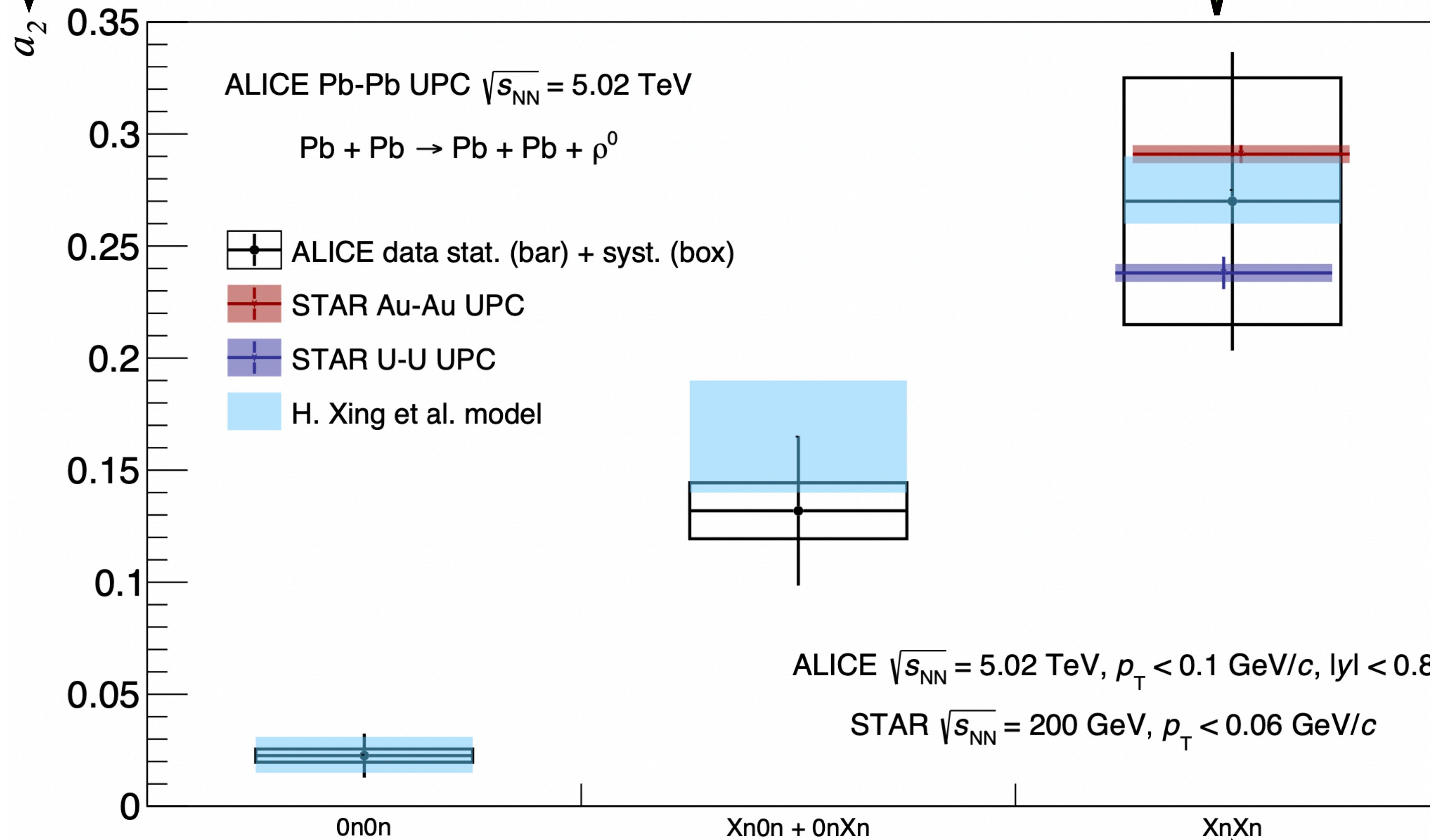
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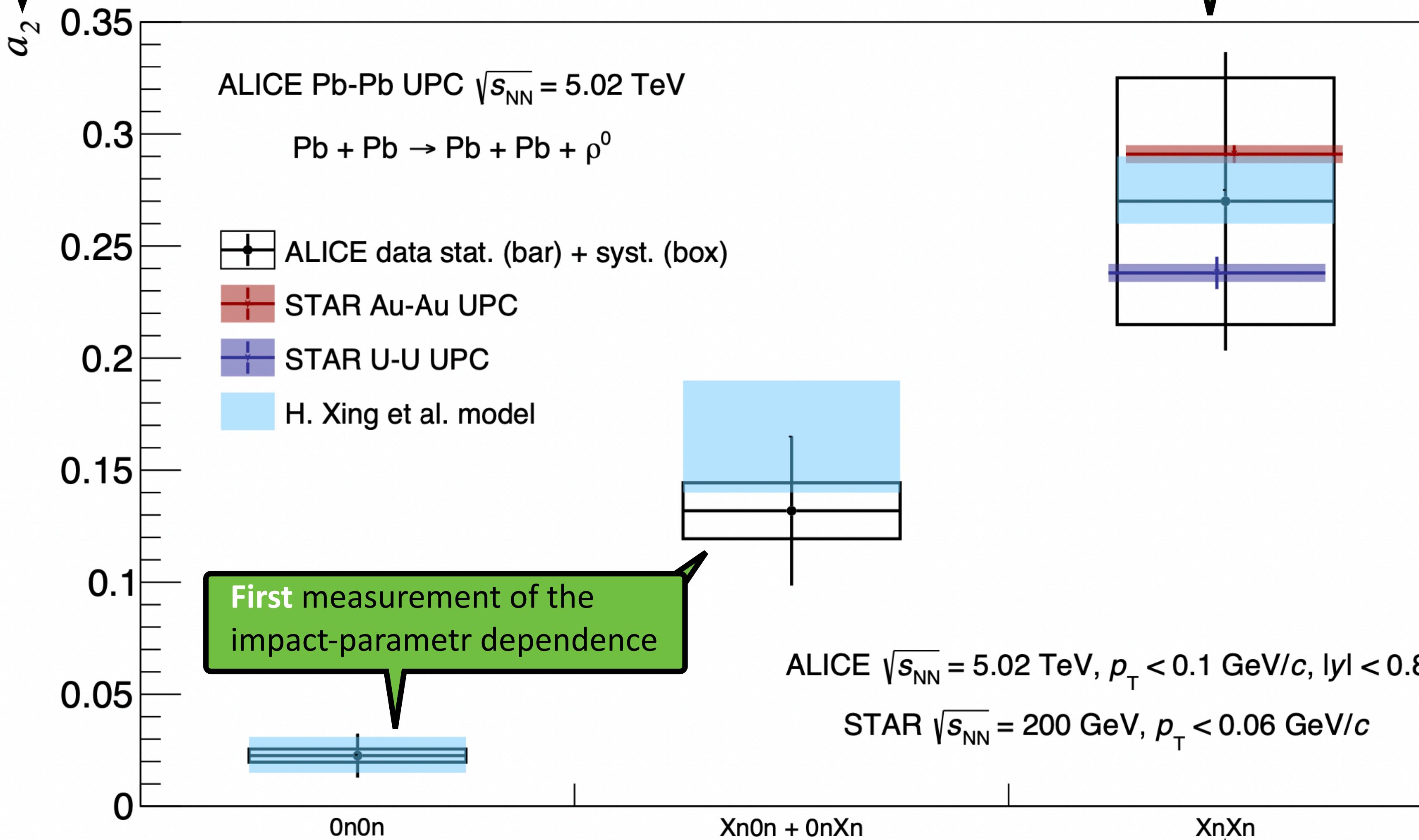


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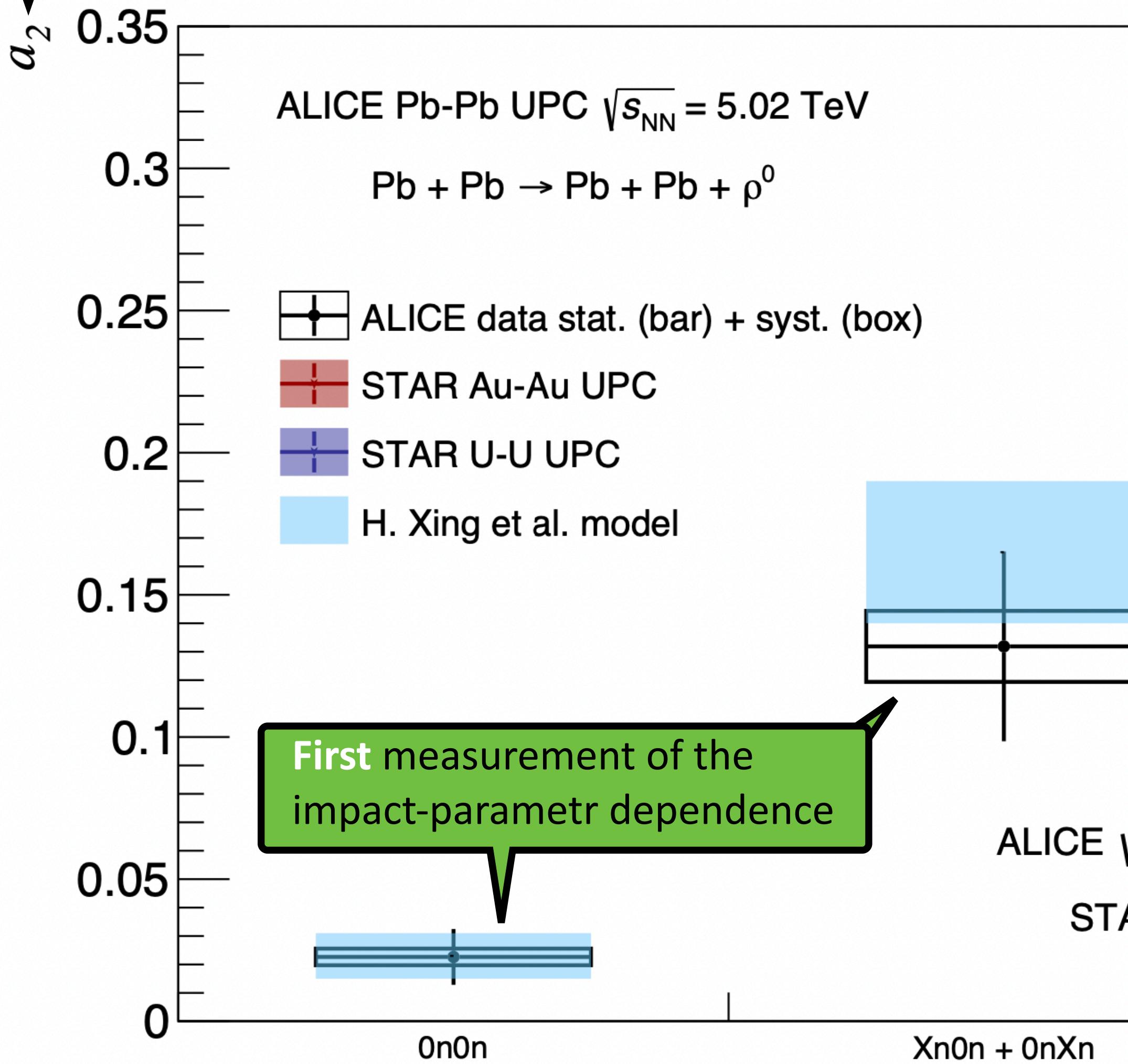
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ALICE and STAR data agree in XnXn

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The prediction describes the data

Neutron classes  $\approx$  impact parameter ranges

# Summary and outlook

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We should have much more (50k vs billions) data in Run 3 :)